

# ERS-2 Wind Scatterometer Cyclic Report

From 24<sup>th</sup> September 2007 to 29<sup>th</sup> October 2007 Cycle 130



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# 1 Introduction and Summary

The document includes a summary of the daily quality control made within the DPQC (Data Processing Quality Control) and various sections describing the results of the investigations and studies of "open-problems" related to the Scatterometer. In each section results are shown from the beginning of the mission in order to see the evolution and to outline possible "seasonal" effects. An explanation for the major events which have impacted the performance since launch is given, and comments about the recent events which occurred during the last cycle are included.

This report covers the period from 24<sup>th</sup> September to 29<sup>th</sup> October (cycle 130) and includes the results of the monitoring activity performed by ESRIN and ECMWF. This document is available on line at: <u>http://earth.esa.int/pcs/ers/scatt/reports/pcs\_cyclic/</u>

#### **Mission events**

The following bullets summarize the major mission facts for cycle 130:

- The ERS-2 satellite was piloted in ZGM throughout the cycle.
- The ESACA processor worked nominally without faults.
- No anomalies occurred on the AMI instrument in the reported period.
- A planned manoeuvre (FCM) has been performed on 18<sup>th</sup> October. During the manoeuvre data accuracy could be degraded. The user can filter out that data set by checking the Doppler and yaw quality flag inside the UWI product or the combined Kp-Yaw flag for the BUFR product.
- Since October, 18th 2007 a new ERS ground station is operative in Chetumal (Mexico) and Scatterometer data is distributed to the users.
- Problem affecting the Hobart station antenna was solved and since 2<sup>nd</sup> October 2007 data was acquired and distributed again.
- Data from all the ground stations was received and distributed to the end users with delay on 26<sup>th</sup> September due to an Esrin hardware failure.
- Missing data from Singapore station from 25<sup>th</sup> to 2<sup>nd</sup> October due to a ground station hardware failure.
- No data was acquired in all ground stations (expect in Kiruna) on 11<sup>th</sup> October from 03 a.m. to 9a.m. due to auxiliary files not disseminated to the ground stations caused by a ground segment problem.
- Missing data from Miami station from 12<sup>th</sup> to 15<sup>th</sup> October due to a ground station problem.
- No data was acquired in all ground stations on 17<sup>th</sup> October from 7 a.m. to 2p.m. UTC due to auxiliary files disseminated with delay to the ground stations.



- From 24<sup>th</sup> October to 29<sup>th</sup> October 2007 meteo files missing or delivered with delay to the ground stations due to a ground segment problem. This caused the ingestion of the wrong meteo table in the processing.
- Missing data from Gatineau station from 26<sup>th</sup> October to 28<sup>th</sup> October due to a network problem.
- For the entire period of cycle 130, ERS-2 Scatterometer data was used in the 4D-Var data assimilation system at ECMWF.

News on the ERS mission is available on line: http://earth.esa.int/ers/new\_ers\_news.html

#### **Data Coverage**

After the on board tape recorder failure in July 2003, data is acquired in real time whenever within the visibility range of a ground station. For cycle 130 data coverage was further extended with the dissemination of data from the Chetumal ground station in Mexico. From 1<sup>st</sup> October data from Hobart station was received again. The data coverage now includes: the North-Atlantic, the Mediterranean, the Caribbean, the Gulf of Mexico, a small part of the Pacific west from the US Canada and Central America, the Chinese and Japanese Sea, a small part of the Indian Ocean South-east of Thailand and Indonesia, the Southern Ocean close to Antarctica and South of Australia and New Zealand.

#### Yaw performance

The result of the monitoring for cycle 130 is an average (per orbit) yaw error angle within the expected nominal range (+/-2 degrees) for most of the orbit.

#### **Calibration performance**

- Calibration data from Transponder are not available since January 2005. This is due to a hardware failure on the transponder. The repair of such device is still under evaluation. The calibration data acquired until 2005 in the ZGM will be re-processed with TOSCA (Tool for Scatterometer Calibration) and the results will be provided in this report when available.
- Due to the regional mission scenario the calibration performances over the Brazilian rain forest are not available because that area was not covered by the ESA ground station. Since 18<sup>th</sup> October a new ground station in Chetumal (Mexico) is operative. The station visibility



now includes the reference test area over the Brazilian rain forest used to monitor the Scatterometer calibration since the beginning of the ERS-2 mission. The calibration monitoring activities has been recovered and will be available from the next cycle.

• The Ocean Calibration monitoring is performed by ECMWF. The average backscatter bias level has improved compared to cycle 129 and is similar to those for a similar period one year ago (see Cyclic Report 120). The gap between the fore/aft and mid beam is still large (especially for the ascending tracks). Average bias levels are less negative (-0.58 dB, was - 0.75), being around 0.20 dB more negative than for nominal data in 2000 (around -0.4 dB). Long-term variations correlate with the yearly cycle, which, given the non-global coverage, is understandable. Therefore, the method of ocean calibration will probably only provide accurate information on calibration levels for globally or yearly data sets.

#### **Instrument performance**

• During the cycle 130 the mean transmitted power evolution had a mean decrease of 0.24 dB per cycle. This value is above the nominal decreasing trend of 0.1 dB/Cycle detected since the beginning of the mission.

• The evolution of the noise power during the cycle 130 was stable. The daily average for the Fore and Aft beam noise is around 1.7 ADC (I) and around 1.6 ADC (Q) respectively. For the Mid beam the noise is not measurable.

• During the cycle 130 the Doppler compensation evolution was stable. The daily average of the CoG of the compensated received signal was around 20 Hz and -20 Hz for the Fore and Aft antenna respectively. For the Mid antenna it was around 220 Hz. The standard deviation of the CoG was around 1500 Hz for the Fore and Aft antenna and around 2750 Hz for the Mid antenna. These values are within the nominal range.

#### **Timeliness performance**

Timeliness performances stayed stable during the cycle 130. Kiruna data are delivered in about 30 min; for the other stations the delivery delay is ranging between 40 and 50 minutes.

#### **Product performance**

During Cycle 130 data was received between 21:01 UTC 24 September 2007 and 20:26 UTC 29 October 2007. Received data was grouped into 6-hourly batches (centred around 00, 06,



12 and 18 UTC). No data was received for the batches of 06 UTC 11 October 2007.

Compared to Cycle 129, the UWI wind speed relative to ECMWF first-guess (FG) fields showed a slightly higher standard deviation (1.46 m/s, was 1.39 m/s). Bias levels were basically unaltered (on average -0.91 m/s, was -0.92 m/s).

The PCS geophysical monitoring reports a wind speed bias (UWI vs 18 or 24 hour forecast) of 0.7 m/s and a speed bias standard deviation around 1.8 m/s The direction deviation performance is stable with more than 98% of the nodes with a wind direction in agreement with the ECMWF winds.

# 2 <u>Calibration Performances</u>

The calibration performances are estimated using three types of target: a man made target (the transponder) and two natural targets (the rain forest and the ocean). This approach allow us to design the correct calibration using a punctual but accurate information from transponders and an extended but noisy information from rain forest and ocean for which the main component of the variance comes from the geophysical evolution of the natural target and from the backscattering models used. These aspects are in the calibration performance monitoring philosophy. The major goals of the calibration monitoring activities are the achievement of a "flat" antenna pattern profile and the assurance of a stable absolute calibration level.

#### 2.1 Gain Constant over transponder

One gain constant is computed per transponder per beam from the actual and simulated twodimensional echo power, which is given as a function of the orbit time and range time. This parameter clearly indicates the difference between "real instrument" and the mathematic model. In order to acquire data over the transponder the Scatterometer must be set in an appropriate operational mode defined as "Calibration Mode". Since January 2001 with the operations in Zero Gyro Mode (ZGM) the satellite attitude is not stable as it was in the nominal Yaw Steering Mode (YSM). In particular there is a non-predictable variation of the yaw error angle along the orbit. For that reason the gain constant data computed by the CALPROC processor, that assumes a stable orbit, are meaningless and a new calibration processor is under development. In the mean time, data from the Transponder are still acquired and archived for future re-processing. The reprocessed gain constants will be provided in this section when available. For the gain constant computed during the nominal



YSM please refer to the Scatterometer cyclic report cycle 60.

#### 2.2 Ocean Calibration

The average sigma0 bias levels (compared to simulated sigma0's based on ECMWF model FG winds) stratified with respect to antenna beam, ascending or descending track and as function of incidence angle (i.e. across-node number) is displayed in Figure 1.

Inter-node and inter-beam dependencies are stable compared to Cycle 129 and similar to that for a similar period one year ago (see Cyclic report 120). The gap between the fore/aft and mid beam is large (especially for the ascending tracks). Average bias levels are less negative (-0.58 dB, was -0.75 dB) being around 0.20 dB more negative than for nominal data in 2000 (around -0.4 dB; see Figure 11 of the reports for Cycle 48 to 59).

Long-term variations correlate with the yearly cycle, which, given the non-global coverage, is understandable. Therefore, the method of ocean calibration will probably only provide accurate information on calibration levels for globally or yearly averaged data sets.

The data volume of descending tracks was about 21% lower than for ascending tracks.



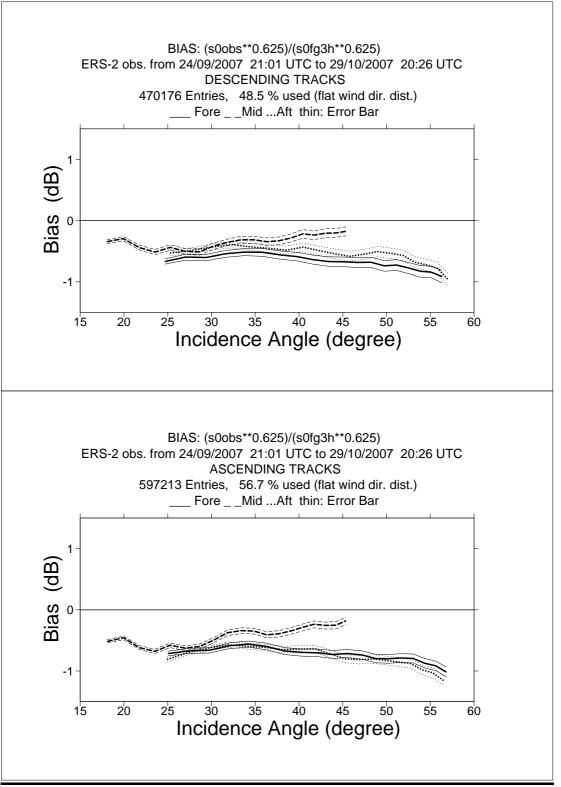


FIGURE 1 ERS-2 Scatterometer Ocean Calibration cycle 130. Ratio of <sigma\_0^0.625>/<CMOD4(First Guess)^0.625> converted in dB for the fore beam (solid line), mid beam (dashed line) an aft beam (dotted line), as a function of incidence angle for descending and ascending tracks. The thin lines indicate the error bars on the estimated mean. First-guess winds are based on the in time closest (+3h, +6h, +9h, or +12h) T511 forecast field, and are bilinearly interpolated in space.



#### 2.3 Gamma-nought over the Brazilian rain forest

Although the transponders give accurate measurements of the antenna attenuation at particular points of the antenna pattern, they are not adequate for fine tuning across all incidence angles, as there are simply not enough samples. The tropical rain forest in South America has been used as a reference distributed target. The target at the working frequency (C-band) of ERS-2 Scatterometer acts as a very rough surface, and the transmitted signal is equally scattered in all directions (the target is assumed to follow the isotropic approximation). Consequently, for the angle of incidence used by ERS-2 Scatterometer, the normalized backscattering coefficient (sigma nought) will depend solely on the surface effectively seen by the instrument:

$$S^0 = S \bullet \cos \theta$$

With this hypothesis it is possible to define the following formula:

$$\gamma^0 = \frac{\sigma^0}{\cos\theta}$$

Using the above equation, the gamma nought backscattering coefficient over the rain forest is independent of the incident angle, allowing the measurements from each of the three beams to be compared. The test area used by the PCS is located between 2.5 degrees North and 5.0 degrees south in latitude and 60.5 degrees West and 70.0 degrees West in longitude. That area has been not covered by the Regional mission scenario (since cycle 86 onwards) and therefore the calibration monitoring activity over the Brazilian rain forest was suspended because no data are available. A new ground station in Chetumal (Mexico) is operative since 18<sup>th</sup> October. The station visibility includes the reference test area over the Brazilian rain forest used to monitor the Scatterometer calibration since the beginning of the ERS-2 mission. The calibration monitoring activities has been recovered and will be available from the next cycle.

#### 2.4 Antenna pattern: Gamma-nought as a function of elevation angle

Due to the regional mission scenario data over the Brazilian rain forest are not available. For that reason the antenna patterns in function of the elevation angle have not been computed.

#### 2.5 Antenna pattern: Gamma-nought as a function of incidence angle

Due to the regional mission scenario data over the Brazilian rain forest are not available. For



that reason the antenna patterns in function of the incidence angle have not been computed.

#### 2.6 Gamma nought histograms and peak position evolution

As the gamma nought is independent from the incidence angle, the histogram of gamma nought over the rain forest is characterized by a sharp peak. The time-series of the peak position gives some information on the stability of the calibration. This parameter is computed by fitting the histogram with a normal distribution added to a second order polynomial:

$$F(x) = A_0 \cdot \exp\left(-\frac{z^2}{2}\right) + A_3 + A_4 \cdot x + A_5 \cdot x^2$$

where:  $z = \frac{x - A_1}{A_2}$ 

The parameters are computed using a non linear least square method called "gradient expansion". The position of the peak is given by the maximum of the function F(x). The histograms are computed weekly (from Monday to Sunday) for each antenna individually "Fore", "Mid" and "Aft" and for ascending and descending passes with a bin size of 0.02 dB. Due to the regional mission scenario data over the Brazilian rain forest are not available and the histograms have not been computed. For the time series since the beginning of the mission please refer to the Scatterometer cyclic report cycle 86.

#### 2.7 Gamma nought image of the reference area

Due to the regional mission scenario data over the Brazilian rain forest are not available and the histograms have not been computed.

#### 2.8 Sigma nought evolution

Due to the regional mission scenario data over the Brazilian rain forest are not available. For that reason none update has been done to the sigma nought evolution time series. For the time series since the beginning of the mission until June 2003 please refer to the Scatterometer cyclic report cycle 86.

#### 2.9 Antenna temperature evolution over the Rain Forest

Due to the regional mission scenario data over the Brazilian rain forest are not available. For the time series since the beginning of the mission please refer to the Scatterometer cyclic report cycle 86.



# 3 Instrument performance

The instrument status is checked by monitoring the following parameters:

• Centre of Gravity (CoG) and standard deviation of the received signal spectrum after the on-ground Doppler Compensation filter. This parameter is useful for the monitoring of the orbit stability, the performances of the Doppler compensation filter, the behavior of the yaw steering mode and the performances of the devices in charge for the satellite attitude (e.g. gyroscopes, Earth sensor, Sun sensor).

- Noise power I and Q channel.
- Internal calibration pulse power.

The latter is an important parameter to monitor the transmitter and receiver chain, the evolution of pulse generator, the High Power Amplifier (HPA), the Traveling Wave Tube (TWT) and the receiver. These parameters are extracted daily from the UWI products and averaged. The evolution of each parameter is characterized by a least square line fit. The coefficients of the line fit are printed in each plot.

#### 3.1 Centre of gravity and standard deviation of received power spectrum

The Figure 2 shows the evolution of the two parameters for each beam since the beginning of the ERS-2 mission and Figure 3 shows the same evolution only for the cycle 130.

The tendency during the nominal Yaw Steering Mode (YSM) period (beginning of the mission since the operation with the Mono Gyro (MGM) Attitude On-board Control System (AOCS) configuration on 7<sup>th</sup> February 2000) is a small and regular increase of the Centre of gravity (CoG) of received spectrum for the three antennae. During the YSM, two small changes can be detected in the CoG evolution. The first change is from 24<sup>th</sup>, January 1996 to 14<sup>th</sup>, March 1996, the second one is from 14<sup>th</sup> February 1997 to 22<sup>nd</sup> April 1997. The reason was a change in the pointing subsystem (DES reconfiguration) side B instead of side A after a depointing anomaly (see table 1 for the list of the all AOCS depointing anomaly occurred during the ERS-2 mission). During these periods side B was switched on. It is important to note that during the first time a clear difference in the CoG of the received spectrum is present only for the Fore antenna (an increase of roughly 100 Hz) while during the second time the change has affected all the three antennae (roughly an increase of 200 Hz, 50 Hz and



50 Hz for the fore, mid and aft antenna respectively).

At the beginning of 2000 the nominal 3-gyroes AOCS configuration (plus one Digital Earth Sensor -DES, and one Digital Sun Sensor -DSS and backups) was no more considered safe because 3 of the six gyros on-board were out of order or very noisy. For that reason the MGM was implemented as default piloting mode. The MGM configuration was designed to pilot the ERS-2 using only one gyro plus the DES and the DSS modules. Scope of ZGM configuration was to extend the satellite lifetime by using the available gyros one at the time.

With the MGM, an increase of roughly 200 Hz was observed at the end of the qualification period. After the AOCS commissioning phase this parameter further evolved within the nominal range with a negligible impact on the data quality.

In MGM configuration, the gyro 5 was used until 7<sup>th</sup> October 2000 when it failed. From 10<sup>th</sup> October 2000 to 24<sup>th</sup> October 2000 the gyro 6 was used. This explains the decrease of roughly 100Hz in the CoG of the received spectrum. From 25<sup>th</sup> October 2000 to 17<sup>th</sup> January 2001 the gyro 1 was used to pilot the ERS-2 satellite. On 17<sup>th</sup> January 2001 the AOCS was upgraded. The new configuration allows piloting the satellite without gyroscopes. Unfortunately a failure of the Digital Earth Sensor (DES A-side) caused ERS-2 to enter in Safe-Mode on the same day. On 25<sup>th</sup> January 2001 gyro #1 also failed.

Satellite attitude was recovered on 5<sup>th</sup> February 2001 with a coarse attitude control mode (EBM). During the period of safe mode the spacecraft had drifted out of the nominal dead band by some 30 Km. The nominal orbit was reached on 6<sup>th</sup> February 2001.

The EBM mode had a strong negative impact on the Scatterometer data quality and the dissemination of data products to end users was discontinued.

After that a series of AOCS upgrades has been implemented in order to improve the satellite attitude: on 30<sup>th</sup> March 2001 the Yaw steering law was re-introduced into the piloting function and on 7<sup>th</sup> June 2001 the Zero Gyro Mode (ZGM) has been implemented as nominal piloting mode. In ZGM the satellite attitude had an improvement in particular for the pitch and yaw error angle. This explains the reduction of the fluctuation in the received signal.

The CoG returns within its nominal value in February 2003 when the new ERS Scatterometer ground processor (ESACA) was put in operation (only for validation purposes) in Kiruna station. ESACA is able to compensate for errors in satellite attitude and to produce calibrated sigma noughts.



The evolution of the standard deviation of the CoG of the received spectrum was stable during the YSM phase. Small peaks are related with the events listed in Table 2. In MGM the evolution was within the nominal range while for the initial phase of the ZGM the performance was strong degraded. This because the on-ground Doppler filters was not able to compensate for the satellite degraded attitude. The introduction of the ESACA processor in February 2003 cured the problem.

On 8th December 2006 10:43 p.m. to 9<sup>th</sup> December 2006 07:18 anomaly in the on board Doppler Compensation occurred. That did not impact on the evolution of the CoG because the ESACA ground processor has compensated the receiver signal for the Doppler frequency shift. The Scat Team has carried out a deep analysis of the anomaly (see the technical note OSME-DPQC-SEDA-TN-06-0328 for further details).

Start of the anoma	Start of the anomaly			End of the anomaly		
24 <sup>th</sup> January	1996	9:10 a.m.	26 <sup>th</sup> January	1996	6:53 p.m.	AOCS depointing anomaly
14 <sup>th</sup> February	1997	1:25 a.m.	15 <sup>th</sup> February	1997	3:44 p.m.	AOCS depointing anomaly
3 <sup>rd</sup> June	1998	2:43 p.m.	6 <sup>th</sup> June	1998	12:47 a.m.	AOCS depointing anomaly
1 <sup>st</sup> September	1999	8:50 a.m.	2 <sup>nd</sup> September	1999	1:28 a.m.	
7 <sup>th</sup> October	2000	4:38 p.m.	10 <sup>th</sup> October	2000	4:49 p.m	depointing anomaly gyro 5 failure
24 <sup>th</sup> October	2000	4:05 p.m.	25 <sup>th</sup> October	2000	12:05 p.m.	depointing anomaly gyro 6 failure
17 <sup>th</sup> January	2001		5 <sup>th</sup> February	2001		gyro 1 failure Satellite in safe mode

 TABLE 1 ERS-2 Scatterometer AOCS depointing anomaly list

 TABLE 2 ERS-2 Scatterometer anomalies in the Doppler Compensation monitoring

Year	Date stop	Year	Reason
1996	27 <sup>th</sup> September	1996	Missing on-board Doppler coefficient
	-		(after cal. DC converter test period) No Yaw Steering Mode
1998	7 <sup>th</sup> June	1998	(after depointing anomaly)
1998	3 <sup>rd</sup> December	1998	Missing on-board Doppler coefficients
1770	5 December	1770	(after AMI anomaly number 228)
2000	17 <sup>th</sup> February	2000	Fine Pointing Mode (FPM)
			(due to AOCS mono-gyro qualification period)
2000	14 <sup>th</sup> April	2000	Fine Pointing Mode (FPM)
	1996         1998         1998         2000	199627th September19987th June19983td December200017th February	1996       27 <sup>th</sup> September       1996         1998       7 <sup>th</sup> June       1998         1998       3 <sup>rd</sup> December       1998         2000       17 <sup>th</sup> February       2000



5 <sup>th</sup> July	2000	5 <sup>th</sup> July	2000	Fine Pointing Mode (FPM) after instrument switch-on
27 <sup>th</sup> September	2000	27 <sup>th</sup> September	2000	Fine Pointing Mode (FPM) to upload AOCS software
Ĩ		Ĩ		patch
2 <sup>nd</sup> November	2000	2 <sup>nd</sup> November	2000	Fine Pointing Mode (FPM)
5 <sup>th</sup> December	2000	6 <sup>th</sup> December	2000	Fine Pointing Mode (FPM) due to orbital manoeuvre
6 <sup>th</sup> February	2001	30 <sup>th</sup> March	2001	Extra Backup Mode (EBM) coarse attitude control
30 <sup>th</sup> March	2001	17 <sup>th</sup> June	2001	ZGM-EBM coarse attitude control
17 <sup>th</sup> June	2001	21 <sup>st</sup> August	2003	ZGM phase. Error in yaw angle not corrected in the
				ground segment processor. Data shall be reprocessed
				with ESACA.
24 <sup>th</sup> March	2004	24 <sup>th</sup> March	2004	Fine Pointing Mode (FPM) due to orbital manoeuvre
25 <sup>th</sup> October	2004	27 <sup>th</sup> October	2004	Series of orbital manoeuvres (OCM and FPM)
10 <sup>th</sup> November	2004	11 <sup>th</sup> November	2004	Intense geomagnetic storm
8 <sup>th</sup> March	2005	8 <sup>th</sup> March	2005	orbital manoeuvre (OCM)
11 <sup>th</sup> March	2005	11 <sup>th</sup> March	2005	orbital manoeuvre (FPM)
2 <sup>nd</sup> November	2005	2 <sup>nd</sup> November	2005	orbital manoeuvre (OCM)
1 <sup>st</sup> March	2006	1 <sup>st</sup> March	2006	orbital manoeuvre (OCM)
3 <sup>rd</sup> November	2006	3 <sup>rd</sup> November	2006	orbital manoeuvre (OCM) at 10:07:46
4 <sup>th</sup> November	2006	4 <sup>th</sup> November	2006	orbital manoeuvre (FCM) at 02:56:53 and 04:37:38
8 <sup>th</sup> December	2006	9 <sup>th</sup> December	2006	Missing on-board Doppler coefficients after AMI
				anomaly from 10:43 p.m. to 9 <sup>th</sup> December 2006 07:18
				a.m.
19 <sup>th</sup> December	2006	19 <sup>th</sup> December	2006	orbital manoeuvre (FCM) at 23:06:12
1 <sup>st</sup> February	2007	1 <sup>st</sup> February	2007	orbital manoeuvre (FCM) at 02:53:31
13 <sup>th</sup> February	2007	13 <sup>th</sup> February	2007	orbital manoeuvre (FCM) at 05:00:15 and 06:40:51
14 <sup>th</sup> February	2007	14 <sup>th</sup> February	2007	orbital manoeuvre (OCM) at 09:30:29
26 <sup>th</sup> April	2007	26 <sup>th</sup> April	2007	Orbital manoeuvre (FCM) at 03:12:03
11 <sup>th</sup> May	2007	11 <sup>th</sup> May	2007	Orbital manoeuvre (FCM) at 02:04:10
13 <sup>th</sup> June	2007	13 <sup>th</sup> June	2007	Orbital manoeuvre (FCM) at 03:41:38
10 <sup>th</sup> September	2007	10 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 02:10:29 and 03:51:05
11 <sup>th</sup> September	2007	11 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 10:01:58
12 <sup>th</sup> September	2007	12 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 02:47:55 and 04:28:31
13 <sup>th</sup> September	2007	13 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 05:37:30 and 07:18:16
14 <sup>th</sup> September	2007	14 <sup>th</sup> September	2007	Orbital manoeuvre (OCM) at 10:07:42
15 <sup>th</sup> September	2007	15 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 23:00:51
16 <sup>th</sup> September	2007	16 <sup>th</sup> September	2007	Orbital manoeuvre (FCM) at 00:41:27
18 <sup>th</sup> October	2007	18 <sup>th</sup> October	2007	Orbital manoeuvre (FCM) at 02:00:00



The Doppler compensation evolution for cycle 130 is showed in Figure 3. The monitoring shows a daily average of the CoG of the compensated received signal around 20 Hz and -20 Hz for the Fore and Aft antenna respectively. For the Mid antenna it was around 230 Hz. The standard deviation of the CoG was around 1500 Hz for the Fore and Aft antenna and around 2750 Hz for the Mid antenna. Those values are within the nominal range.



#### **ERS-2** WindScatterometer: **DOPPLER** COMPENSATION Evolution (UWI)

Least-square poly. fit fore beam Least-square poly. fit mid beam Least-square poly. fit aft beam Center of gravity =  $-22.00 + (0.0120)^*$ day Standard Deviation =  $5329.2 + (-0.927)^*$ day Center of gravity =  $-716.9 + (0.2598)^*$ day Standard Deviation =  $5974.8 + (-0.793)^*$ day Center of gravity =  $-280.9 + (0.0820)^*$ day Standard Deviation =  $5475.7 + (-0.962)^*$ day

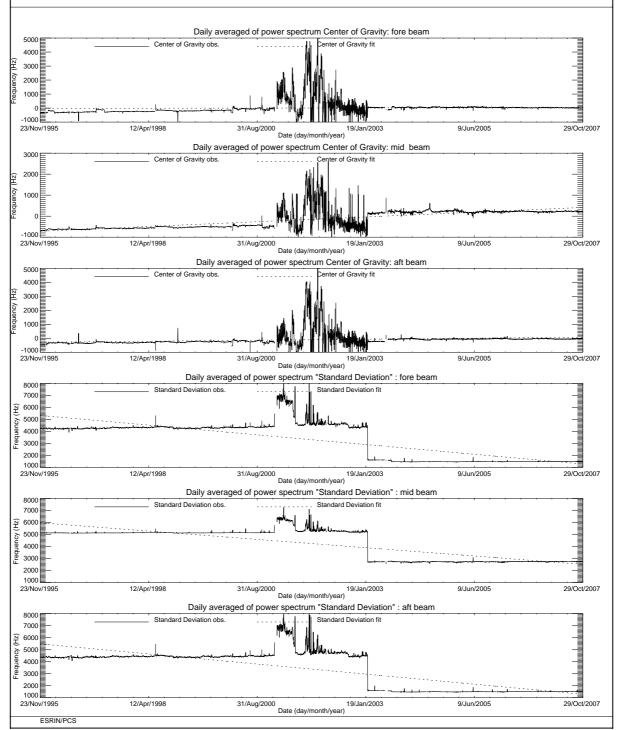


FIGURE 2 ERS-2 Scatterometer: Centre of Gravity and standard deviation of received power spectrum since the beginning of the mission.



#### **ERS-2 WindScatterometer: DOPPLER COMPENSATION Evolution (UWI)**

Least-square poly. fit fore beam Least-square poly. fit mid beam Least-square poly. fit aft beam

Center of gravity =  $11.784 + (0.0213)^*$ day Standard Deviation =  $1488.8 + (0.0882)^*$ day Center of gravity =  $216.75 + (0.0300)^*$ day Standard Deviation =  $2739.5 + (0.2405)^*$ day Center of gravity =  $-24.91 + (0.4415)^*$ day Standard Deviation =  $1497.6 + (0.1133)^*$ day

		averaged of power spectrum Center of Gravity: fore bea	m <u> </u>	
400	Center of Gravity ob	s. Center of Gravity fit		
<sup>[</sup> 귀 200				
Z200				~~~~
-200 -				
-400				
24/Sep/2007	1/Oct/2007	8/Oct/2007 Date (day/month/year) 15/Oct/2007	22/Oct/2007	29/Oct/2007
<u> </u>	<u> </u>	averaged of power spectrum Center of Gravity: mid bea	m	
400	Center of Gravity ob	s. Center of Gravity fit		
(ÎH) 200 E			······································	
(TH) 200				
-200 -				
-400				
24/Sep/2007	1/Oct/2007	8/Oct/2007 15/Oct/2007 Date (day/month/year)	22/Oct/2007	29/Oct/2007
· · · · · ·		averaged of power spectrum Center of Gravity: aft bear	n	,
400	Center of Gravity ob	s. Center of Gravity fit		
(F) 200				
(TH) 200 L L L L L L L L L L L L L L L L L L				
e -200				
-400				
24/Sep/2007	1/Oct/2007	8/Oct/2007 15/Oct/2007 Date (day/month/year)	22/Oct/2007	29/Oct/2007
4000 =		eraged of power spectrum "Standard Deviation" : fore be	eam	=
	Standard Deviation of	bbs Standard Deviation fit		
H)				
(H) (2) (2) (2) (2) (2) (2) (2) (2				
1000 H				_
۰ <u>ا</u>				
24/Sep/2007	1/Oct/2007	8/Oct/2007 15/Oct/2007 Date (day/month/year)	22/Oct/2007	29/Oct/2007
4000 =	Daily av	reraged of power spectrum "Standard Deviation" : mid be	eam	
	Standard Deviation of	bbs. Standard Deviation fit		
Ê <sup>3000</sup>				
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o				
24/Sep/2007	1/Oct/2007	8/Oct/2007 15/Oct/2007 Date (day/month/year)	22/Oct/2007	29/Oct/2007
4000 =		veraged of power spectrum "Standard Deviation" : aft be	am	
	Standard Deviation of	bbs. Standard Deviation fit		
Î Î 3000				
2000 2000				
(7) 2000				
I E				
0 <u>⊨</u> 24/Sep/2007	1/Oct/2007	8/Oct/2007 15/Oct/2007 Date (day/month/year)	22/Oct/2007	 29/Oct/2007
ESRIN/PCS		Date (day/month/year)		

FIGURE 3 ERS-2 Scatterometer: Centre of Gravity and standard deviation of received power spectrum for cycle 130.



#### 3.2 Noise power level I and Q channel

The results of the monitoring are shown in Figure 4 (long-term) and Figure 5 (cycle 130). The first set of three plots presents the noise power evolution for the I channel while the second set shows the Q channel. From the plots one can see that the noise level is more stable in the I channel than in the Q one. The I and Q receivers are inside the same box and any external interference should affect both channel. The fact that the receivers are closer to the ATSR-GOME electronics could have some impact but there is no clear explanation on that behavior. From 5<sup>th</sup> December 1997 until November 1998 some high peaks appear in the plots. These high values for the daily mean are due to the presence for these special days of a single UWI product with an unrealistic value in the noise power field of its Specific Product Header. The analysis of the raw data used to generate these products lead in all cases to the presence of one source packet with a corrupted value in the noise field stored into the source packet Secondary Header. The reason why noise field corruption is beginning from 5<sup>th</sup> December 1997 and last until November 1998 is at present unknown. It is interesting to note that at the beginning of December 1997, we started to get as well the corruption of the Satellite Binary Times (SBTs) stored in the EWIC product. The impact in the fast delivery products was the production of blank products starting from the corrupted EWIC until the end of the scheduled stop time. A change in the ground station processing in March 1998 overcame this problem.

Since 9<sup>th</sup> August 1998 until March 2000 some periods with a clear small instability in the noise power have been recognized, Table 3 gives the detailed list.

Start date	Stop date	Year	
9 <sup>th</sup> August	26 <sup>th</sup> October	1998	
29 <sup>th</sup> November	6 <sup>th</sup> December	1998	
23 <sup>rd</sup> December	24 <sup>th</sup> December	1998	
7 <sup>th</sup> June	10 <sup>th</sup> June	1999	
17 <sup>th</sup> August	22 <sup>nd</sup> August	1999	
8 <sup>th</sup> September	9 <sup>th</sup> September	1999	
3 <sup>rd</sup> October	8 <sup>th</sup> October	1999	
16 <sup>th</sup> October	18 <sup>th</sup> October	1999	
26 <sup>th</sup> October	28 <sup>th</sup> October	1999	
25 <sup>th</sup> December	2 <sup>nd</sup> January	2000	
10 <sup>th</sup> February	11 <sup>th</sup> February	2000	
19 <sup>th</sup> March	26 <sup>th</sup> March	2000	

 TABLE 3 ERS-2 Periods with instability in the noise power



To better understand the instability of the noise power the PCS has carried out investigations in the Scatterometer raw data (EWIC) to compute the noise power with more resolution. The result is that for the orbits affected by the instability the noise power had a decrease of roughly 0.7 dB for the fore and aft signals and a decrease of roughly 0.6 dB in the mid beam case (see the report for the cycle 42). The decrease of the noise power during the orbits affected by the instability is comparable with the decrease of the internal calibration level that occurred during the same orbits. The reason of this instability (linked to the AMI anomalies) is still unknown. On 28<sup>th</sup> February 2003 the Scatterometer receiver gain has been increased by 3 dB to optimize the usage of the on-board ADC converter. This explains the increase of the noise for the Fore and Aft beam channel. For the mid beam channel the noise still remains not measurable.

On 17<sup>th</sup> February 2006 a high peak was detected in the noise power, causing the daily average for that day very high. The case has been deeply investigated and a technical note (Ref OSME-DPQC-SEDA-TN-06-0163) is available. The cause was an acquisition problem that corrupted one source packet and not an instrument anomaly. The same happened on April 24<sup>th</sup> 2006 (cycle 115).

On 8<sup>th</sup> September 2006 a high peak in the noise power of the Mid beam has been detected. The event occurred between 17:41:54 and 17:42:43 (UTC) and the noise power reached the value of 43 ADC (fore beam) and 19 ADC (mid beam). Those values had affected the daily average and are clear present in the plots of the Figure 4. That anomaly has been deeply investigated in the Technical Note OSME-DPQC-SEDA-TN-06-0251 and cannot be linked to any anomaly in the acquired data. The conclusion of the investigation was that a problem had occurred in the transmitter or in the pulse generator of the AMI instrument. At that time the AMI was in wind only mode so no additional comparison with SAR data can be done. Similar peaks had been noted also for September 15<sup>th</sup> and 18<sup>th</sup>. ESOC has checked the Mission Plan and noticed that in all three events the peak in the noise power occurred very close to 6 minutes after the start of a Wind mode and 40 minutes after ascending node crossing.

The evolution of the noise power during the cycle 130 was stable. The daily average for the Fore and Aft beam noise is around 1.7 ADC (I) and around 1.6 ADC (Q) respectively. For the Mid beam the noise is not measurable.



#### **ERS-2 WindScatterometer: NOISE Level Evolution (UWI)** I = 735.21 +(0.2472)\*day Q = 684.78 +(0.2341)\*day Least-square line fit fore beam: I channel: No line fit standard deviation too hight Q channel: No line fit standard deviation too hight Least-square line fit aft beam: I = 734.11 +(0.2374)\*day Q channel: No line fit standard deviation too hight Channel I Fore Beam: daily averaged (min = 466.700 max = 4208.80 mean = 1275.07 std = 369.357 2000 Noise power fit Noise power obs 1500 ADC 1000 0.001 500 ot 23/Nov/1995 12/Apr/1998 31/Aug/2000 19/Jan/2003 9/Jun/2005 29/Oct/2007 Date (day/month/year) Channel I Mid Beam: daily averaged (min = -0.300000 max = 9045.50 mean = 13.1107 std = 323.328) 2000 Noise power fit e power obs mhuhuhu 1500 Chit Q 1000 0.001 500 0 31/Aug/2000 Date (day/month/year) 23/Nov/1995 9/Jun/2005 29/Oct/2007 19/Jan/2003 12/Apr/1998 Channel I Aft Beam: daily averaged (min = 410.700 max = 9861.60 mean = 1257.84 std = 411.900) 2000 Noise pow Noise nower fit 1500 Jnit ADC 1000 1111 0.001 500 Ξ ٥ŀ 23/Nov/1995 12/Apr/1998 9/Jun/2005 29/Oct/2007 31/Aug/2000 19/Jan/2003 Date (day/month/year) Channel Q Fore Beam: daily averaged (min = 392.000 max = 4434.70 mean = 351.625 1196.32 std 2000 Noise power obs se nower fit 1500 Ë ADC 1000 0.001 500 0 23/Nov/1995 12/Apr/1998 31/Aug/2000 19/Jan/2003 9/Jun/2005 29/Oct/2007 Date (dav/month/vear) Channel Q Mid Beam: daily averaged (min = -0.300000 max = 8299.80 mean = 6.97943 std = 222.980) 2000 ..... Noise power fit minimu 1500 Ë Q 1000 0.00 500 0 23/Nov/1995 12/Apr/1998 31/Aug/2000 19/Jan/2003 9/Jun/2005 29/Oct/2007 Date (day/month/year) Channel Q Aft Beam: daily averaged (min = 330.400 max = 17371.8 mean = 1174.83 std = 581.603) 2000 1500 , ⊔ ADC 1000 0.001 500 23/Nov/1995 12/Apr/1998 9/Jun/2005 29/Oct/2007 31/Aug/2000 19/Jan/2003 Date (day/month/year) ESRIN/PCS

FIGURE 4 ERS-2 Scatterometer: noise power I and Q channel since the beginning of the mission.



#### **ERS-2** WindScatterometer: NOISE Level Evolution (UWI)

Least-square line fit fore beam: Least-square line fit mid beam: Least-square line fit aft beam:  $I = 1713.4 + (-0.417)^* day$ 

I = 1758.7 +(-0.658)\*day I = 0.0968 +(0.0013)\*day Q = 1660.0 +(-0.830)\*day  $Q = 0.1605 + (-0.002)^* day$  $Q = 1603.0 + (-0.530)^* day$ 

2000	Noise power obs.		. Noise power fit		· · · · ]
1500					
E					E
000					
E					E
500					
٥ <u> </u>					
4/Sep/2007	1/Oct/2007		15/Oct/2007 month/year)	22/Oct/2007	29/Oct/2
.000	Noise power obs.	ily averaged (min = 0.0000	00 max = 1.00000 mean = 0.	144444 std = 0.197765)	· · · · ¬
Ē	Noise power obs.		. Noise power in		E
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000					
E					=
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0E					=
4/Sep/2007	1/Oct/2007	8/Oct/2007 Date (day/	15/Oct/2007 month/year)	22/Oct/2007	29/Oct/2
2000	Channel I Aft Beam: c	laily averaged (min = 1681.	50 max = 1725.00 mean = 17	706.09 std = 11.0644)	<u> </u>
	Noise power obs.		Noise power fit		
500					
E					=
000					
500					_
E					E
0 <u> </u>	1/Oct/2007	8/Oct/2007	15/Oct/2007	22/Oct/2007	
4/360/2007	1/06/2007	Date (day/	month/year)	22/06/2007	25/00/2
2000	Channel Q Fore Beam:	daily averaged (min = 161	5.10 max = 1667.80 mean = 1	1645.54 std = 15.0238)	
-	Noise power obs.		. Noise power fit		
	Noise power obs.		Noise power fit		
500	Noise power obs.		Noise power fit		
1500	Noise power obs.		. Noise power fit		
1500	Noise power obs.		. Noise power fit		
	Noise power obs.		. Noise power fit		
1500		8/0+/2007		22/0ct/2007	
	Noise power obs.	8/Oct/2007 Date (day/	15/Oct/2007	22/Oct/2007	29/Oct/2
500	1/Oct/2007	Date (day/	15/Oct/2007		29/Oct/2
500	1/Oct/2007	Date (day/	15/Oct/2007 month/year)		29/Oct/2
500 500 500 0	1/Oct/2007 Channel Q Mid Beam: da	Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
500	1/Oct/2007 Channel Q Mid Beam: da	Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
1500	1/Oct/2007 Channel Q Mid Beam: da	Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
1500 1000 500 4/Sep/2007 2000 1500 1000 1	1/Oct/2007 Channel Q Mid Beam: da	Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
1500	1/Oct/2007 Channel Q Mid Beam: da	Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
1500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs.	Date (day	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0. . Noise power fit	147222 std = 0.206309)	
500 500 500 500 0 0 500 0 500	1/Oct/2007 Channel Q Mid Beam: da	Date (day/ aily averaged (min = 0.0000	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0.		29/Oct/2
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs.	Date (day/ aily averaged (min = 0.0000 8/Oct/2007 Date (day/	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007	147222 std = 0.206309) 22/Oct/2007	
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs.	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) )00 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year)	147222 std = 0.206309) 22/Oct/2007	
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	
1500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	
500	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	
	1/Oct/2007 Channel Q Mid Beam: da Noise power obs. 1/Oct/2007 Channel Q Aft Beam: d	Date (day) aily averaged (min = 0.0000 8/Oct/2007 Date (day) daily averaged (min = 1564	15/Oct/2007 month/year) 000 max = 1.00000 mean = 0. . Noise power fit 15/Oct/2007 month/year) .90 max = 1622.50 mean = 1	147222 std = 0.206309) 22/Oct/2007	

FIGURE 5 ERS-2 Scatterometer: noise power I and Q channel for cycle 130.



#### **3.3** Power level of internal calibration pulse

For the internal calibration level, the results are shown in Figure 6 (long-term) and Figure 7 (cycle 130). The high value of the variance in the fore beam until August, 12<sup>th</sup> 1996 is due to the ground processing. In fact all the blank source packets ingested by the processor were recognized as Fore beam source packets with a default value for the internal calibration level. The default value was applicable for ERS-1 and therefore was not appropriate for ERS-2 data processing. On August 12<sup>th</sup>, 1996 a change in the ground processing LUT overcame the problem. Since the beginning of the mission a power decrease is detected. The power decrease is regular and affects the AMI when it is working in wind-only mode, wind/wave mode and image mode indifferently. The average power decrease is around 0.08 dB per cycle (0.0022 dB/day) and is clearer after August, 6<sup>th</sup> 1996 when the calibration subsystem has been changed. The reason of the power decrease is because the TWT is not working in saturation, so that a variation in the input signal is visible in the output. The variability of the input signal can be two-fold: the evolution of the pulse generator or the tendency of the switches between the pulse generator and the TWT to reset themselves into a nominal position. These switches were set into an intermediate position in order to put into operation the Scatterometer instrument (on 16<sup>th</sup> November 1995). To compensate for this decrease, on 26<sup>th</sup> October 1998 (cycle 37) 2.0 dB were added to the Scatterometer transmitted power and on 4<sup>th</sup> September 2002 (cycle 77) were added 3.0 dB. On 28<sup>th</sup> February 2003 (cycle 82) the Scatterometer receiver gain was increased by 3 dB to improve the usage of the on-board ADC converter. These events are clearly displayed by the large steps show in Figure 6. Since 9<sup>th</sup> August 1998 until March 2000 the internal calibration level shows instability after an AMI or platform anomaly (see reports from cycle 35 to cycle 52). This instability is very well correlated with the fluctuations observed in the noise power. On 13th July 2000 a high peak (+3.5 dB) was detected in the transmitted power. This event has been investigated

deeply by PCS and ESOC. The results of the analysis are reported in the technical note "ERS-2 Scatterometer: high peak in the calibration level" available in the PCS. The high transmitted power was detected after an arcing event which occurred inside the HPA. After that event the transmitted power had an average increase of roughly 0.14 dB.

During the cycle 130 the mean transmitted power evolution had a mean decrease of 0.24 dB. This value is higher than the nominal decreasing trend of 0.1 dB/Cycle detected since the beginning of the mission.



#### ERS-2 WindScatterometer: Internal CALIBRATION Level Evolution (UWI)

Least-square polynomial fit fore beam Least-square polynomial fit mid beam Least-square polynomial fit aft beam gain (dB) per day 0.0000 gain (dB) per day 0.0000 gain (dB) per day 0.0000 1055.37 +(-0.00159177)\*day 312.895 +(-0.000907928)\*day 1043.28 +(-0.000755096)\*day

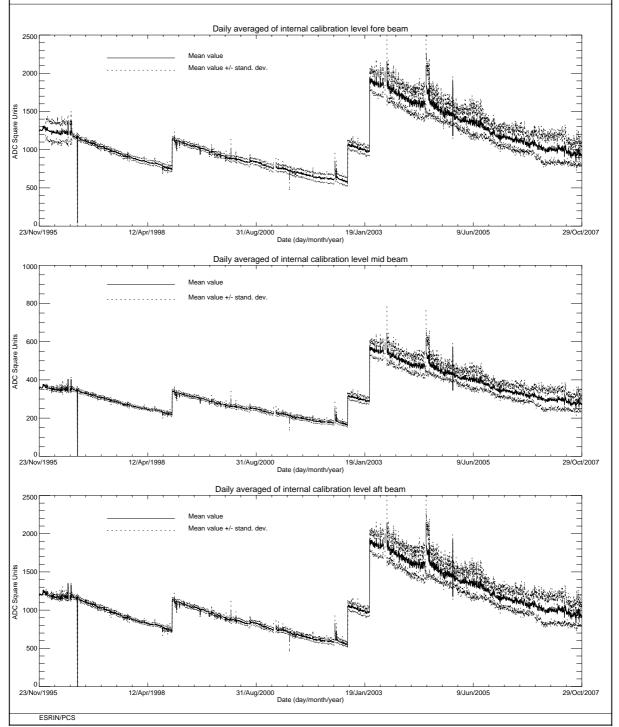


FIGURE 6 ERS-2 Scatterometer: power of internal calibration pulse since the beginning of the mission.



#### ERS-2 WindScatterometer: Internal CALIBRATION Level Evolution (UWI)

Least-square polynomial fit fore beam Least-square polynomial fit mid beam Least-square polynomial fit aft beam gain (dB) per day -0.0072 gain (dB) per day -0.0069 gain (dB) per day -0.0075 975.812 +(-1.57537)\*day 287.574 +(-0.445103)\*day 972.125 +(-1.62283)\*day

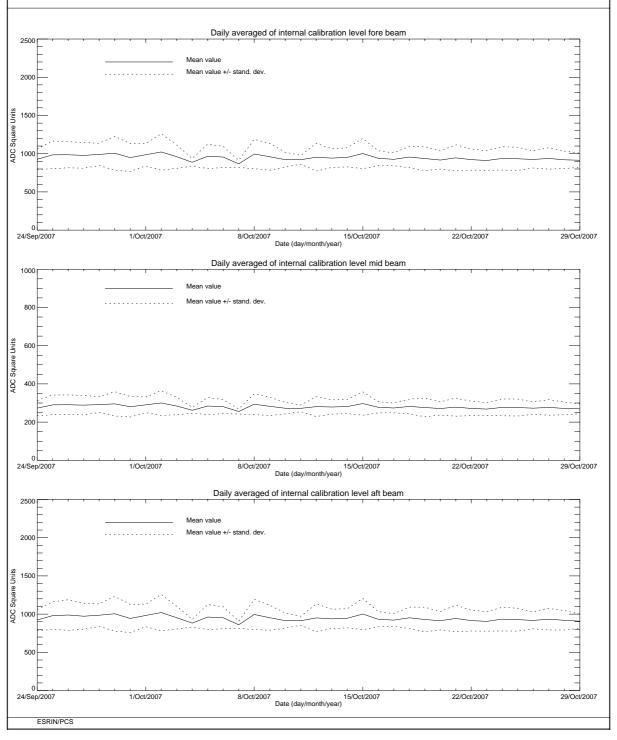


FIGURE 7 ERS-2 Scatterometer: power of internal calibration level cycle 130.



## 4 Products performance

The PCS carries out a quality control of the winds generated from the WSCATT data. External contributions to this quality control (from ECMWF) are also reported in this chapter.

#### 4.1 **Products availability**

One of the most important points in the monitoring of the products performance is their availability. The Scatterometer is a part of ERS payload and it is combined with a Synthetic Aperture Radar (SAR) into a single Active Microwave Instrument (AMI). The SAR users requirements and the constraints imposed by the on-board hardware (e.g. amount of data that can be recorded in the on-board tape) set rules in the mission operation plan.

The principal rules that affected the Scatterometer instrument data coverage are:

• Over the Ocean the AMI is in wind/wave mode (Scatterometer with small SAR imagettes acquired every 30 sec.) and the ATSR-2 is in low rate data mode.

• Over the Land the AMI is in wind only mode (only Scatterometer) and the ATSR-2 is in high rate mode. (Due to on board recorder capacity, ATSR-2 in high rate is not compatible with SAR wave imagettes acquisitions.) This strategy preserves the Ocean mission.

• The SAR images are planned as consequence of users' request.

Moreover:

- since July 16<sup>th</sup> 2003 the ERS-2 Low Rate mission is continued within only the visibility of ESA ground stations over Europe, North Atlantic, the Arctic and western North America. The reason was the failure of both on-board tape recorders.
- During the cycles 64 92 (June 2001 since 25<sup>th</sup> February 2004) the AMI instrument was operated in wind/wave mode also over the land. The reason was because the SAR wave data was used to estimate the satellite mispointing along the full orbit. Since 25<sup>th</sup> February onwards the nominal mission scenario has been resumed, with the AMI instrument in wind only mode over the land (and consequently ATSR was operated again in High Rate over land). The mispointing performances (in particular the yaw error angle) along the full orbit are computing by analyzing the Scatterometer data.

In order to maximize the data coverage, after the on-board tape recorder failure, an upgrade of the ERS ground segment acquisition scenario has been performed. In that framework the following has been implemented:



- Since September 7<sup>th</sup> 2003 the ground station in Maspalomas, Gatineau and Prince Albert are acquiring and processing data for all the ERS-2 satellite passes within the station visibility (apart from passes for which other satellites have an higher priority).
- To further increase the wind coverage of the North Atlantic area, since December 8<sup>th</sup>, 2003 is operative a new ground Station in West Freugh (UK) and data from this new station are available to the user since mid January 2004. Due to its location, the West Freugh acquisitions have some overlap with those from three other ESA stations, Kiruna, Gatineau or Maspalomas. The station overlap depends on the relative orbit of the satellite. Consequentially, overlapping wind Scatterometer LBR data may be included in two products. Since the two products are generated at different ground stations the overlap may not be completely precise, with a displacement up to 12 Km and slight differences in the wind data itself.
- Since March, 3<sup>rd</sup> 2004, Matera station is acquiring and processing low rate bit data for all the passes for which is planned a SAR acquisition. This means for the Scatterometer data coverage a limited improvement due to the fact that is acquired only a passage with some planned SAR activity.
- Since February 2005 a new acquisition station in Miami (US) is in operations. This new station allows a full data coverage of the Gulf of Mexico and part of the Pacific Ocean on the west Mexico coast.
- Since 25<sup>th</sup>, June 2005 a new acquisition stations have been put into operations in Beijing. It covers part of China and Oriental Asia.
- Since 5<sup>th</sup> July 2005 McMurdo ground station is operational in the South Pole. It covers all the Antarctic region.
- Since 5<sup>th</sup> December 2005 the Hobart station is operational and it is covering the Australian and New Zealand area. Hobart data has been disseminated into BUFR format since February 13<sup>th</sup> 2006.
- At the end of August 2006 a new ground station in Singapore has been installed and products are distributed to the users since October 19<sup>th</sup> 2006.

Figure 8 shows the AMI operational modes for cycle 130. Each segment of the orbit has different color depending on the instrument mode: brown for wind only mode, blue for wind-wave mode and green for image mode. The red and yellow colors correspond to gap modes (no data acquired). For cycle 130 the percentage of the ERS-2 AMI activity is shown in table 4. The value for cycle 130 shows an increase of SAR activity at descending passes with respect to the cycle 129 (14.69%, was 11.75%).



Ami Mode	Ascending passes	Descending passes	
Wind and Wind-Wave	88.35 %	74.75%	
Image	4.35 %	14.69 %	
Gap and others	7.29 %	10.54 %	

 TABLE 4 ERS-2 AMI activity (cycle 130)

Table 5 reports the major data lost (day or more) due to the test periods, AMI and satellite anomalies or ground segment anomalies occurred after 6<sup>th</sup> August, 1996 (before that day for many times data were not acquired due to the DC converter failure).

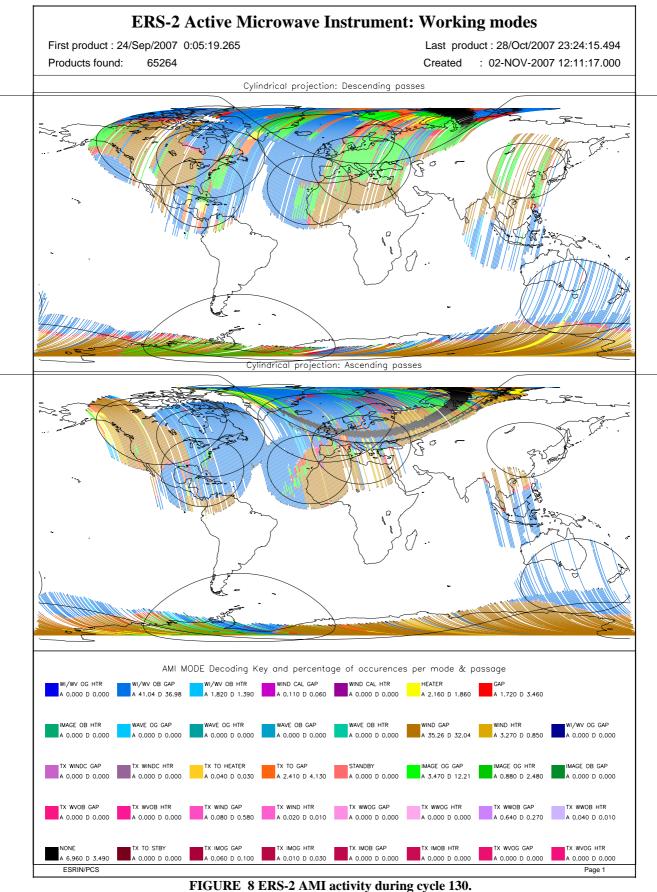
Start date	Stop Date	Reason
September 23 <sup>rd</sup> , 1996	September 26 <sup>th</sup> , 1996	ERS 2 switched off due to a test period
February 14 <sup>th</sup> , 1997	February 15 <sup>th</sup> , 1997	ERS 2 switched off due to a depointing anomaly
June 3 <sup>rd</sup> , 1998	June 6 <sup>th</sup> , 1998	ERS 2 switched off due to a depointing anomaly
November 17 <sup>th</sup> , 1998	November 18 <sup>th</sup> , 1998	ERS 2 switched off to face out Leonide meteor storm
September 22 <sup>nd</sup> 1999	September 23 <sup>rd</sup> 1999	ERS 2 switched off due to Year 2000 certification test
November 17 <sup>th</sup> , 1999	November 18 <sup>th</sup> , 1999	ERS 2 switched off to face out Leonide meteor storm
December 31 <sup>st</sup> ,1999	January 2 <sup>nd</sup> , 2000	ERS 2 switched off Y2K transition operation
February 7 <sup>th</sup> ,2000	February 9 <sup>th</sup> , 2000	ERS 2 switched off due to new AOCS s/w up link
June 30 <sup>th</sup> , 2000	July 5 <sup>th</sup> , 2000	ERS 2 Payload switched off after RA anomaly
July 10 <sup>th</sup> , 2000	July 11 <sup>th</sup> , 2000	ERS 2 Payload reconfiguration
October 7 <sup>th</sup> , 2000	October 10 <sup>th</sup> 2000	ERS 2 Payload switched off after AOCS anomaly
January 17th , 2001	February 5 <sup>th</sup> , 2001	ERS 2 Payload switched off due to AOCS anomaly
May 22 <sup>nd</sup> , 2001	May 24 <sup>th</sup> , 2001	ERS 2 Payload switched off due to platform anomaly
May 25 <sup>th</sup> , 2001	May 25 <sup>th</sup> , 2001	AMI switched off due thermal analysis
November 17 <sup>th</sup> , 2001	November 18 <sup>th</sup> , 2001	ERS 2 switched off to face out Leonide meteor storm
November 27 <sup>th</sup> , 2001	November 28 <sup>th</sup> , 2001	ERS 2 payload off due to 1Gyro Coarse Mode
		commissioning
March 8 <sup>th</sup> , 2002	March 20 <sup>th</sup> , 2002	ERS 2 payload unavailability after RA anomaly
May 19 <sup>th</sup> ,2002	May 24 <sup>th</sup> 2002	AMI switched off due to arc events
May 24 <sup>th</sup> , 2002	May 28 <sup>th</sup> , 2002	AMI partially switched off due to arc events
May 31 <sup>st</sup> 2002	June 3 <sup>rd</sup> 2002	Gatineau orbits partially acquired due to antenna problem
June 4 <sup>th</sup> , 2002	June 5 <sup>th</sup> , 2002	AMI partially switched-off due to arc events
July 25 <sup>th</sup> , 2002	July 25 <sup>th</sup> , 2002	AMI switched off HPA voltage too low
September 11 <sup>th</sup> , 2002	September 11 <sup>th</sup> , 2002	AMI switched off macrocommand transfer error

 TABLE 5 ERS-2 Scatterometer mission major data lost (day or more) after 6<sup>th</sup>, August 1996



November 17 <sup>th</sup> , 2002	November 18 <sup>th</sup> , 2002	ERS-2 switched off to face out Leonide meteor storm
December 9 <sup>th</sup> , 2002	December 10 <sup>th</sup> , 2002	IDHT anomaly no data recorded on board
December 20 <sup>th</sup> , 2002	December 20 <sup>th</sup> 2002	IDHT anomaly no data recorded on board
January 14 <sup>th</sup> , 2003	January 14 <sup>th</sup> , 2003	IDHT anomaly no data recorded on board
May 6 <sup>th</sup> , 2003	May 19 <sup>th</sup> , 2003	AMI off due to bus reconfiguration
June 22 <sup>nd</sup> , 2003	July 16 <sup>th</sup> ,2003	IDHT recorders test no data acquired
Since July 16 <sup>th</sup> ,2003		Regional Mission Scenario. Data available only within the visibility of ESA ground station
May 21 <sup>st</sup> , 2004	May 25 <sup>th</sup> , 2004	AMI in refuse mode due to excessive HPA arcing
June 22 <sup>nd</sup> ,2004	June 22 <sup>nd</sup> , 2004	AMI in refuse mode due to excessive HPA arcing
September 23 <sup>rd</sup> , 2004	September 24 <sup>th</sup> , 2004	AMI switched down
December 16 <sup>th</sup> , 2004	December 17 <sup>th</sup> , 2004	AMI memory test
December 26 <sup>th</sup> , 2004	December 26 <sup>th</sup> , 2004	IDHT anomaly. No data acquired
December 27 <sup>th</sup> , 2004	December 28 <sup>th</sup> , 2004	Payload off due to on board anomaly
January 23 <sup>rd</sup> , 2005	January 23 <sup>rd</sup> , 2005	AMI switched down (00.51 a.m. – 1.26 p.m.)
February 26 <sup>th</sup> , 2005	February 26 <sup>th</sup> , 2005	AMI switched down (01.20 a.m 12.37 a.m.)
May 23 <sup>rd</sup> , 2005	May 24 <sup>th</sup> , 2005	ERS 2 payload unavailability after RA anomaly
Jun 20 <sup>th</sup> , 2005	Jun 21 <sup>st</sup> , 2005	AMI switched off caused by RBI status error (08:44 p.m. – 10:13 a.m.)
December 8 <sup>th</sup> , 2006	December 8 <sup>th</sup> , 2006	AMI switched down to Standby/MCMD Execution Inhibited due to Format Acquisition Error (02:04 p.m. – 10:43 p.m.)
April, 13 <sup>th</sup> , 2007	April 13 <sup>th</sup> , 2007	AMI Switched down to Standby/MCMD Execution Inhibited due to Format Acquisition Error (03:10 a.m. – 12.06 p.m.)
May, 22 <sup>nd</sup> , 2007	May, 22 <sup>nd</sup> , 2007	AMI Switched down to Standby/MCMD Execution Inhibited due to Acquisition Errors (01:50 p.m. – 07.04 p.m.)
June, 10 <sup>th</sup> , 2007	June, 10 <sup>th</sup> , 2007	AMI Switched down to Standby/MCMD Execution Inhibited due to Format Length and ICU Begin Identifier Errors (00:55 a.m. – 10.13 a.m.)
June, 11 <sup>th</sup> , 2007	June, 12 <sup>th</sup> , 2007	AMI Switched down to Standby/MCMD Execution Inhibited due to Emergency Switchdown requested by AMI ICU (10:39 p.m. – 10.49 a.m.)
July, 27 <sup>th</sup> , 2007	July, 27 <sup>th</sup> , 2007	AMI switchdown to Standby/MCMD Execution Inhibited due to RBI Status Error (00:44 a.m 09:43 a.m).







#### 4.2 PCS Geophysical Monitoring

The routine analysis is summarized in the plots of figure 9; from top to bottom:

• the monitoring of the valid sigma-nought triplets per day.

• the evolution of the wind direction quality. The ERS wind direction (for all nodes and only for those nodes where the ambiguity removal has worked properly) is compared with the ECMWF forecast. The plot shows the percentage of nodes for which the difference falls in the range -90.0, +90.0 degrees.

• the monitoring of the percentage of nodes whose ambiguity removal works successfully.

• the comparison of the wind speed deviation: (bias and standard deviation) with the ECMWF forecast.

The results since August 6<sup>th</sup>, 1996 until the beginning of the operation with the Zero Gyro Mode (ZGM) in January 2001 can be summarized as:

• High quality wind products has been distributed since Mid March 1996 (end of calibration and validation phase)

• The number of valid sigma-nought distributed per day was almost stable with a small increase after June 29<sup>th</sup>, 1999 due to the dissemination in fast delivery of the data acquired in the Prince Albert station (Canada).

• The wind direction is very accurate for roughly 93% of the nodes, the ambiguity removal processing successfully worked for more than 90.0% of the nodes.

• The UWI wind speed shows an absolute bias of roughly 0.5 m/s and a standard deviation that ranges from 2.5 m/s to 3.5 m/s with respect to the ECMWF forecast.

• The wind speed bias and its standard deviation have a seasonal pattern due to the different winds distribution between the winter and summer season.

• Two important changes affect the speed bias plot.

• the first is on June 3<sup>rd</sup>, 1996 due to the switch from ERS-1 to ERS-2 data assimilation in the meteorological model.

• the second which occurred at the beginning of September 1997, is due to the new monitoring and assimilation scheme in ECMWF algorithms (4D-Var).

• Since 19<sup>th</sup> April 1999 two set of meteo-table (meteorological forecast centred at 00:00 and 12:00 of each day) are used in the ground processing. This allowed the processing of wind data with 18 and 24 hours meteorological forecast instead of the 18, 24, 30 36 hours forecast. The comparison between data processed with the 18-24 hours forecast instead of 30-36 hours forecast shown an increase in the number of ambiguity removed nodes with a neutral impact in the daily statistics.



• The mono-gyro AOCS configuration (see report for cycle 50) that was operative from 7<sup>th</sup> February 2000 to 17<sup>th</sup> January 2001 did not affect the wind data performance.

During the Zero Gyro Mode (ZGM) phase the dissemination of the fast delivery Scatterometer data to the users has been interrupted on 17<sup>th</sup> January 2001 due to degraded quality in sigma noughts and winds. The satellite attitude in ZGM is slightly degraded and the "old" ground processor was not able to produce calibrated data anymore. For that reason a redesign of the entire ground processing has been carried out and since August 21<sup>st</sup> 2003 the new processor named ERS Scatterometer Attitude Corrected Algorithm (ESACA) is operative in all the ESA ground station and data was redistributed to the user.

Although for a long period data was not distributed, the PCS has monitored the data quality (as shown in Figure 9) and the results during that period can be summarized as:

At the beginning of the ZGM (January 2001 - end July 2001) the number of valid nodes has clear drop from 190000 per day to 9000 per day. This because the satellite attitude was strong degraded and the received signal had a very high Kp figure (in particular for the far range nodes). For the valid nodes, due to no calibrated sigma nought, the quality of the wind was very poor, the distance from the cone was high and the wind speed bias was above 1.5 m/s.

At the end of July 2001 the ZGM has been tuned and the satellite attitude had an improvement. This explains the increase of the number of valid nodes (returned around the nominal level) and the improvements in the wind speed bias (around 0.5 m/s).

On 4<sup>th</sup> February 2003, a beta version of the new ESACA processor has been put in operation in Kiruna for validation and the monitoring of the data quality has been done only for the new ESACA data. The number of valid nodes slight decreased because Kiruna station process only 9 of 14 orbits per day. The wind speed direction deviation had a clear improvement because ESACA implements a new ambiguity removal algorithm (MSC) and the ambiguity removal rate is now stable at 100% (the MSC is able to remove ambiguity for all the nodes). The wind speed bias had a clear drop from 0.5 to -0.5 m/s. That value is closer to the nominal one (around -0.2 m/s). As reported in the previous cyclic reports the beta version of ESACA had some calibration problem for the near range nodes and this explains why the data quality does not match exactly the one obtained in the nominal YSM. That problem has been overcome with the final release of the ESACA processor put into operation on August 21<sup>st</sup> 2003. On June 22<sup>nd</sup> the failure of the on-board tape recorder discontinued the ERS global mission (see section 4.1) and this explains the low number of valid nodes available after that day.

The performances of ESACA winds delivered between August 2003 and September 2004 are affected by land contamination. Around costal zones many Sea nodes have a strong contribution of Land backscattering and the retrieved wind is not correct. An optimization of



the Land/Sea flag in the ground processing has been carried out during the cycle 98. In the statistics computed by PCS on the fast delivered winds the Land contamination has been removed by using a refined Land/Sea mask. Also the ice contamination has been removed with a simple geographical filter. With these new setting the PCS statistics are very similar to the ones reported by ECMWF.

For cycle 130 the wind performances stayed stable. The wind speed bias (UWI vs 18 or 24 hour forecast) was roughly 0.7 m/s and the speed bias standard deviation was around 1.8 m/s. Missing statistics from 24<sup>th</sup> to 28<sup>th</sup> October are due to a ground segment problem that affected the Meteo files dissemination chain. Some meteo tables were not disseminated to the ground station therefore data was processed with wrong meteorological tables.

The reduced number of nodes available on 27<sup>th</sup> September is due to a dissemination problem caused by a hardware failure in Esrin. The reduced number of nodes on 11<sup>th</sup> and 17<sup>th</sup> October is due to a reduced amount of data acquired caused by an auxiliary files dissemination problem to the ground stations.

The wind direction deviation for cycle 130 was good with more than 98% of the nodes wind direction in agreement with the ECMWF forecast.



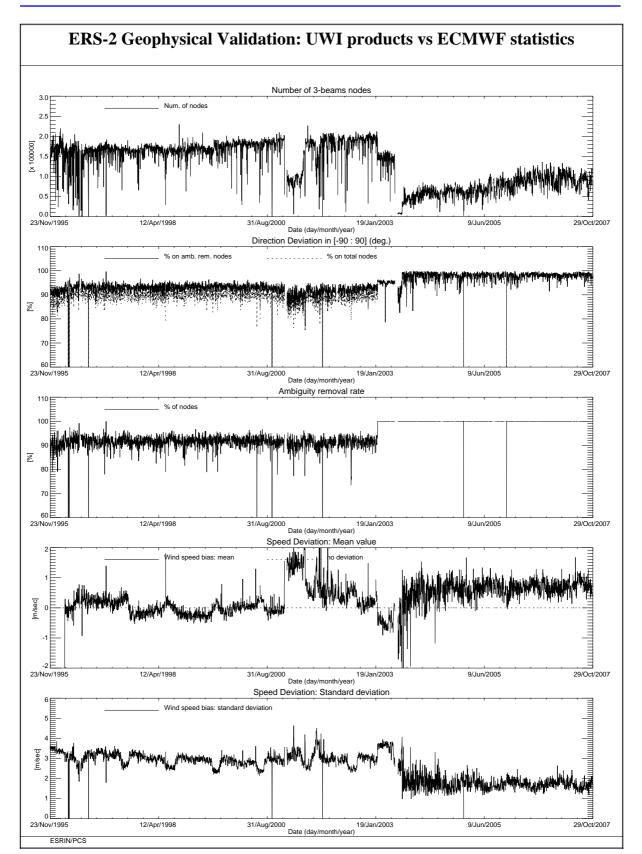


FIGURE 9 ERS-2 Scatterometer: wind products performance since the beginning of the mission.



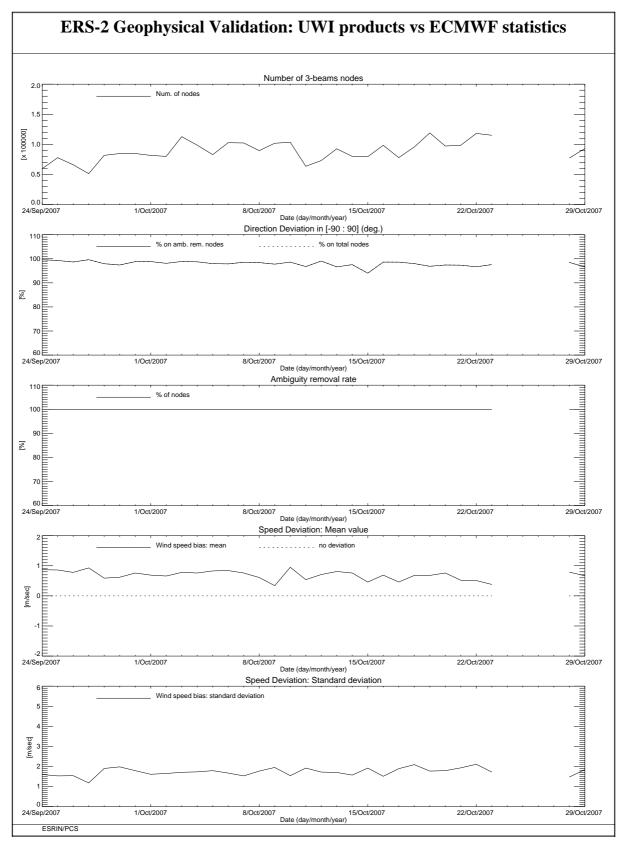


FIGURE 10 ERS-2 Scatterometer: wind products performance for cycle 130.



### 4.3 ECMWF Geophysical Monitoring

The quality of the UWI product was monitored at ECMWF for Cycle 130. Results were compared to those obtained from the previous Cycle, as well for data received during the nominal period in 2000 (up to Cycle 59). No corrections for duplicate observations were applied.

During Cycle 130 data was received between 21:01 UTC 24 September 2007 and 20:26 UTC 29 October 2007. Received data was grouped into 6-hourly batches (centred around 00, 06, 12 and 18 UTC). No data was received for the batches of 06 UTC 11 October 2007 (due to missing data acquisition caused by a ground segment problem).

Data is being recorded whenever within the visibility range of a ground station. From 1 October 2007 onwards data from Hobart station data was received again. On 18 October 2007, first data was received from a new station at Chetumal in Mexico, which extends coverage somewhat more westwards in the East Pacific. Quality of the data obtained from Hobart and Chetumal were found nominal. For cycle 130, data coverage was over the North-Atlantic, the Mediterranean, the Caribbean, the Gulf of Mexico, a small part of the Pacific west from the US, Canada and Central America, the Chinese and Japanese Sea, a small part of the Indian Ocean South-East of Thailand and Indonesia, and the Southern Ocean close to the Antarctica and south of Australia and New Zealand (see Figure 12).

The asymmetry between the fore and aft incidence angles showed several large peaks.

Compared to Cycle 129, the UWI wind speed relative to ECMWF first-guess (FG) fields showed a slightly higher standard deviation (1.46 m/s, was 1.39 m/s). Bias levels were basically unaltered (on average -0.91 m/s, was -0.92 m/s).

The Cycle-averaged evolution of performance relative to ECMWF first-guess (FG) winds is displayed in Figure 11. Figure 12 shows global maps of the over Cycle 130 averaged UWI data coverage and wind climate, Figure 13 for performance relative to FG winds.



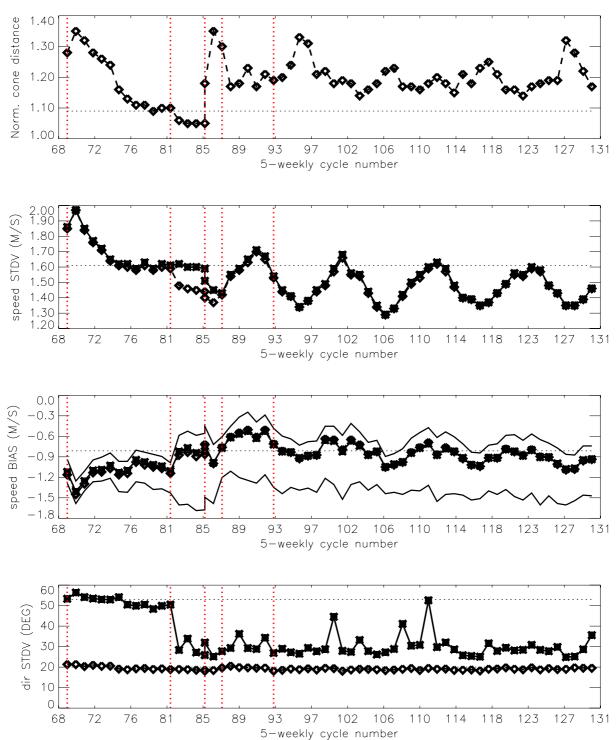
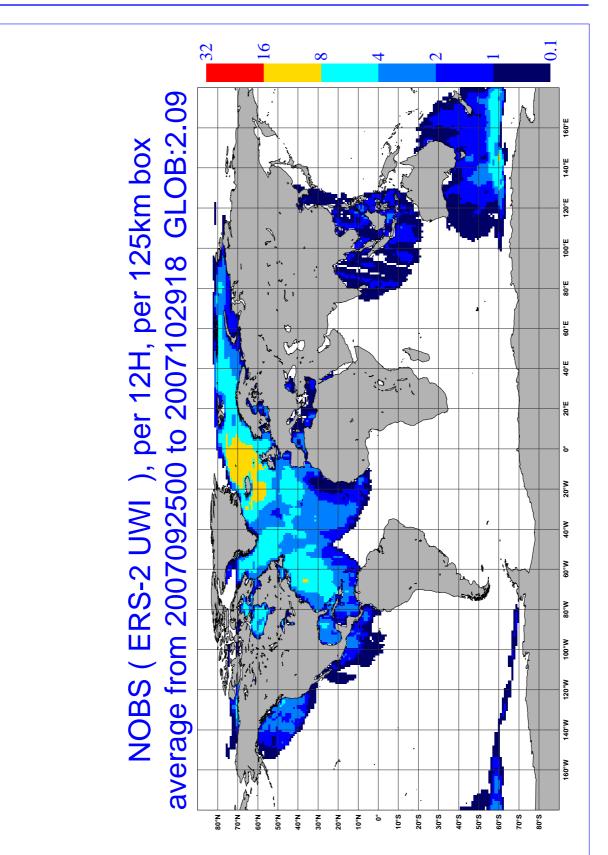


FIGURE 11 Evolution of the performance of the ERS-2 Scatterometer averaged over 5-weekly cycles from 12 December 2001 (cycle 69) to 29 October 2007 (end cycle 130) for the UWI product (solid, star) and de-aliased winds based on CMOD4 (dashed, diamond). Results are based on data that passed the UWI QC flags. For cycle 85 two values are plotted; the first value for the global set, the second one for the regional set. Dotted lines represent values for cycle 59 (5 December 2000 to 17 January 2001),i.e. the last stable cycle of the nominal period. From top to bottom panel are shown the normalized distance to the one (CMOD4 only) the standard deviation of the wind speed compared to FG winds, the corresponding bias (for UWI winds the extreme inter-node averages are shown as well),and the standard deviation of wind direction compared to FG.







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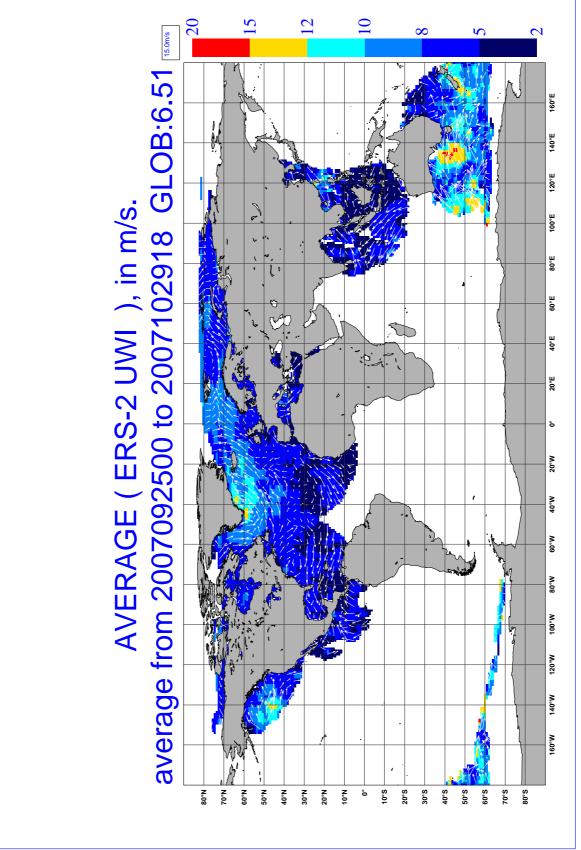
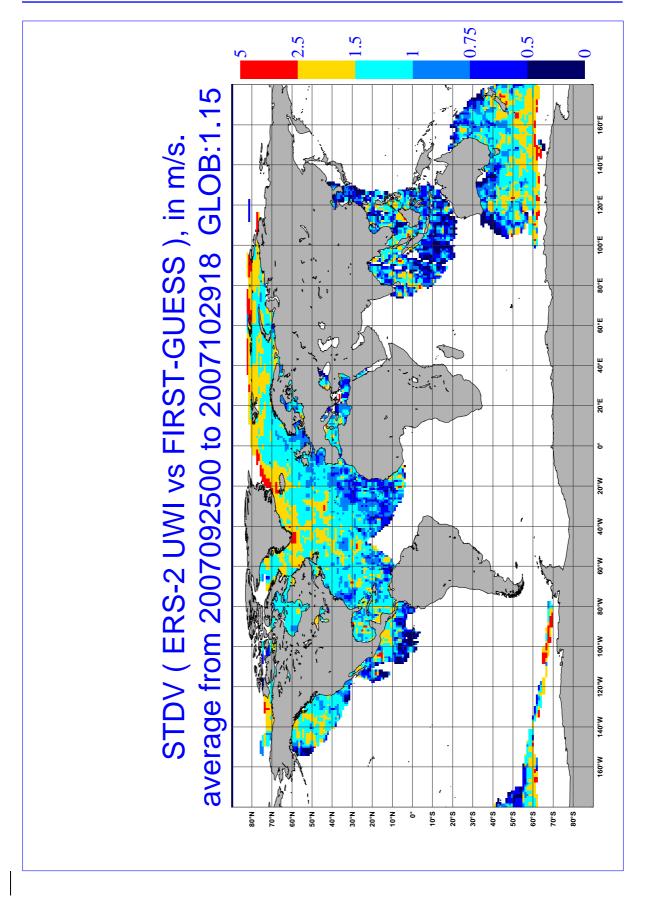


FIGURE 12 Average number of observations per 12H and per 125km grid box (top panel) and wind-climate (lower panel) for UWI winds that passed the UWI flags QC and a check on the collocated ECMWF land and sea-ice mask.







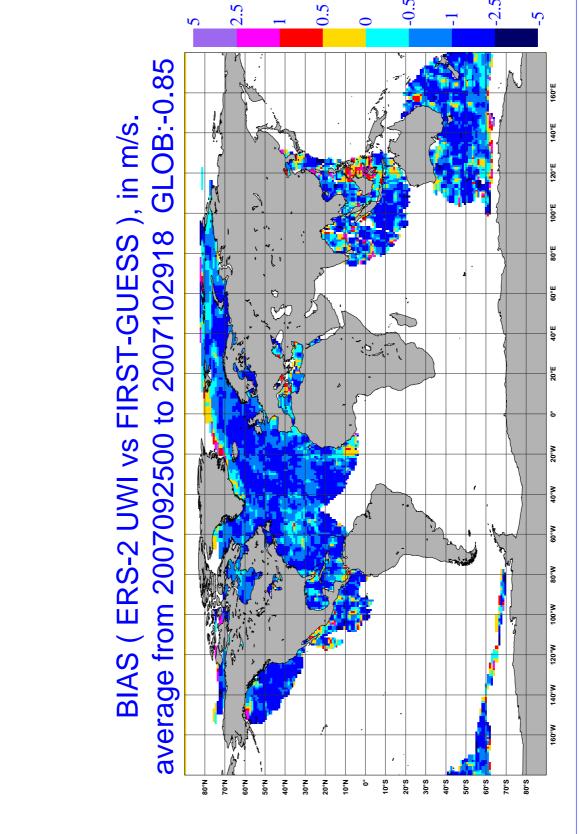


FIGURE 13 The same as Figure 12, but now for the relative bias (top panel) and standard deviation (lower panel) with ECMWF first-guess winds.



### 4.3.1 Distance to cone history

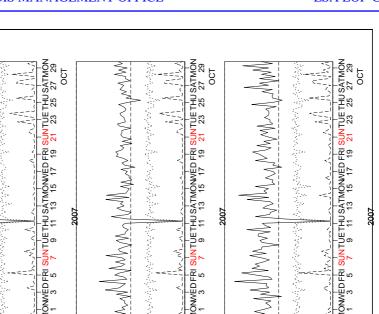
The distance to the cone history is shown in Figure 14. Curves are based on data that passed all QC, including the test on the K\_p-yaw flag, and subject to the land and sea-ice check at ECMWF (see cyclic report 88 for details).

Like for cycle 129, time series are (due to lack of statistics) very noisy, especially for the near-range nodes. Most spikes were found to be the result of low data volumes (such as on 11<sup>th</sup> October for no data received for the batches of 06 UTC due to no data acquired at the ground stations).

Compared to cycle 129, the average level was lower (1.17 versus 1.22), however, higher (by 7%) than for nominal data (see top panel Figure 11).

The fraction of data that did not pass QC is displayed in Figure 14 as well (dash curves).





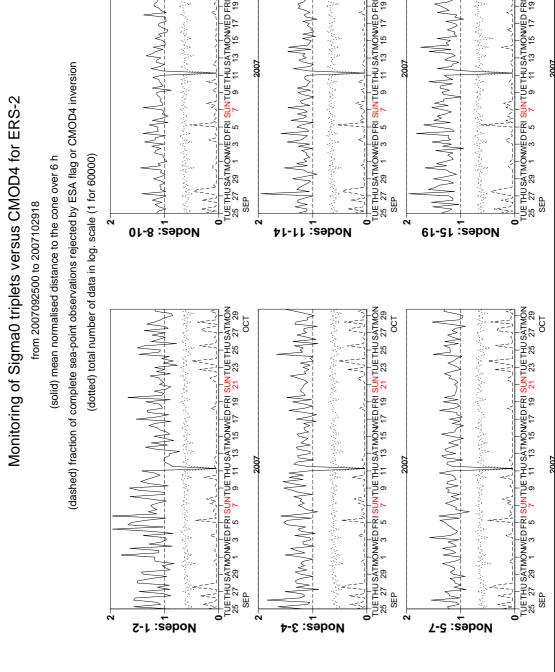


FIGURE 14 Mean normalized distance to the cone computed every 6 hours for nodes 1-2, 3-4, 5-7, 8-10, 11-14 and 15-19 (solid curve close to 1 when no instrumental problems are present). The dotted curve shows the number of incoming triplets in logarithmic scale (1 corresponds to 60,000 triplets) and the dashed one indicates the fraction of complete (based on the land and sea-ice mask at ECMWF) sealocated triplets rejected by ESA flags, or by the wind inversion algorithm (0: all data kept, 1: no data kept).



## 4.3.2 UWI minus First-Guess history

In Figure 15, the UWI minus ECMWF first-guess wind-speed history is plotted. The history plot shows a few peaks, which are usually the result of low data volume.

Figure 19 displays the locations for which UWI winds were more than 8 m/s weaker (top panel) and more than 8 m/s stronger (lower panel) than FG winds. Like for cycle 129, such collocations are isolated, and often indicate meteorologically active regions, for which UWI data and ECMWF model field show reasonably small differences in phase and/or intensity. Deviations near the poles are the result of imperfect sea-ice flagging.

Two examples for a mismatch between ECMWF and UWI winds are the capture of two lowpressure system, one south-east from New Zealand on 13 October 2007 (top panel), and one south off the Mexican coast on 18 October 2007 (lower panel).

The case near New Zealand, which was acquired by Hobart station, shows some differences in the shape of the elongated low between model and scatterometer winds. The UWI winds clearly suffer from some de-aliasing problems.

For the case in the East Pacific, close to Mexico, ERS-2 shows a patch of strong westerly winds with nearly constant wind direction, where the ECMWF model winds have a more circular flow around the pressure minimum.

Average bias levels and standard deviations of UWI winds relative to FG winds are displayed in Table 6. From this it follows that the bias of UWI winds was very similar (-0.91 m/s, was - 0.92 m/s), being around -0.1 m/s more negative than for nominal data in 2000.

	Cycle 129		Cycle 130	
	UWI	CMOD4	UWI	CMOD4
Speed STDV	1.39	1.39	1.46	1.46
Node 1-2	1.40	1.39	1.53	1.51
Node 3-4	1.36	1.35	1.44	1.43
Node 5-7	1.33	1.32	1.40	1.40
Node 8-10	1.35	1.35	1.43	1.42
Node 11-14	1.37	1.37	1.44	1.44
Node 15-19	1.39	1.40	1.44	1.45
Speed BIAS	-0.92	-0.93	-0.91	-0.91

Table 6 Wind speed and direction biases



DATA QUALITY AND ALGORITHMS MANAGEMENT OFFICE

Node 1-2	-1.42	-1.40	-1.43	-1.40
Node 3-4	-1.21	-1.17	-1.18	-1.14
Node 5-7	-1.98	-0.95	-0.94	-0.91
Node 8-10	-0.79	-0.79	-0.76	-0.76
Node 11-14	-0.72	-0.74	-0.72	-0.74
Node 15-19	-0.72	-0.76	-0.73	-0.77
Direction STDV	28.7	19.6	35.6	19.6
Direction BIAS	-2.4	-2.4	-1.4	-1.9

On a longer time scale seasonal bias trends are observed (see Figure 11). As was highlighted in the previous cyclic reports, it is believed that this yearly trend is partly induced by changing local geophysical conditions. Strong indication for this is a similar trend observed for QuikSCAT data when restricted to an area well-covered by ERS-2 (20N-90N, 80W-20E).

Figure 25 shows time series for that area for both ERS-2 (top panel) and QuikSCAT (lower panel) for the period between 1 January 2004 and 29 October 2007 (end of cycle 130). Results are displayed for at ECMWF actively assimilated data, i.e., CMOD5 winds for ERS-2 and 4%-reduced QuikSCAT winds on a 50km resolution. Note the increase in wind speed for ERS-2 since the introduction of the new model cycle at ECMWF on 7 June 2007. It reflects the switch from the CMOD5 to CMOD5.4 model function, which has increased ERS-2 wind speed by 0.48 m/s.

The standard deviation of UWI wind speed versus ECMWF FG was, compared to cycle 129, somewhat higher (1.46 m/s, was 1.39 m/s).

For cycle 130 the (UWI - FG) direction standard deviations were mostly ranging between 20 and 40 degrees (Figure 17), representing nominal variations. However, between 25 and 27 October 2007, the performance was severely reduced. Since no reduction was observed for at ECMWF de-aliased winds, it indicates temporal problems in the de-aliasing procedure in the UWI processing. Averaged over the entire cyclic period, STDV for UWI wind direction has increased (35.6 degrees, was 28.7 degrees), mainly induced by the two-day anomalous period. For at ECMWF de-aliased winds no trend was observed (STDV 19.6, unchanged).

Degradation in the wind direction from  $24^{th}$  to  $28^{th}$  October is due to missing meteo tables in the ground stations caused by a ground segment problem.



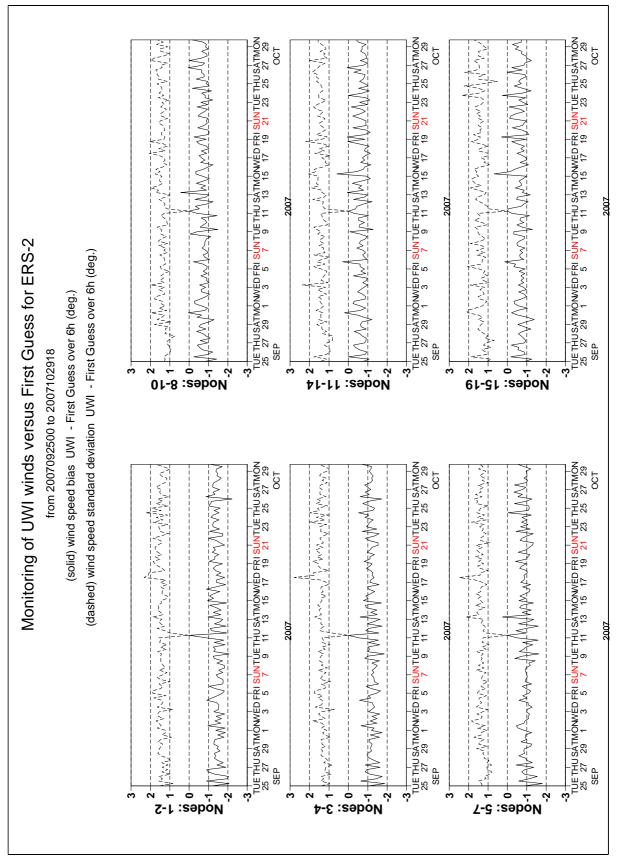


FIGURE 15 Mean (solid line) and standard deviation (dashed line) of the wind speed difference UWI - first guess for the data retained by the quality control.



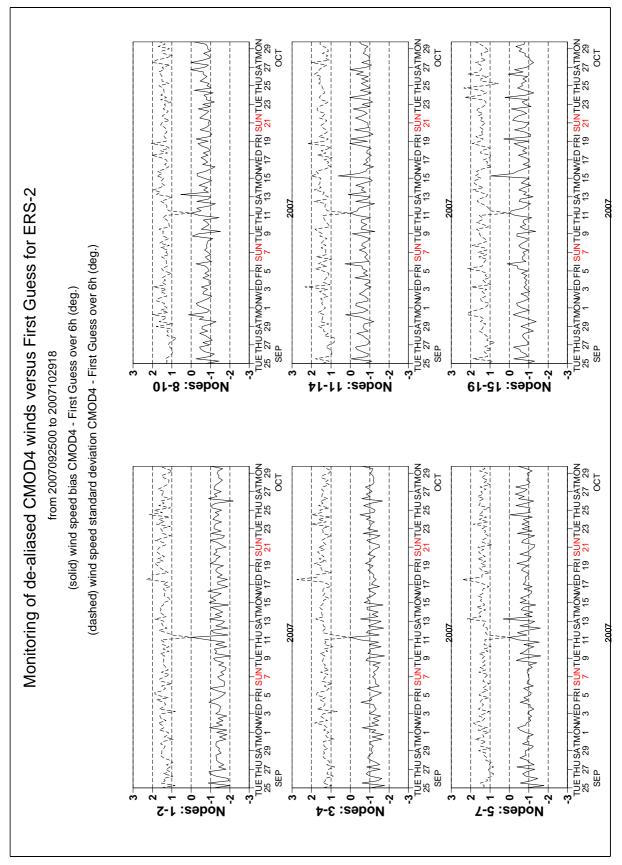


FIGURE 16 Same as Fig. 15, but for the de-aliased CMOD4 data.





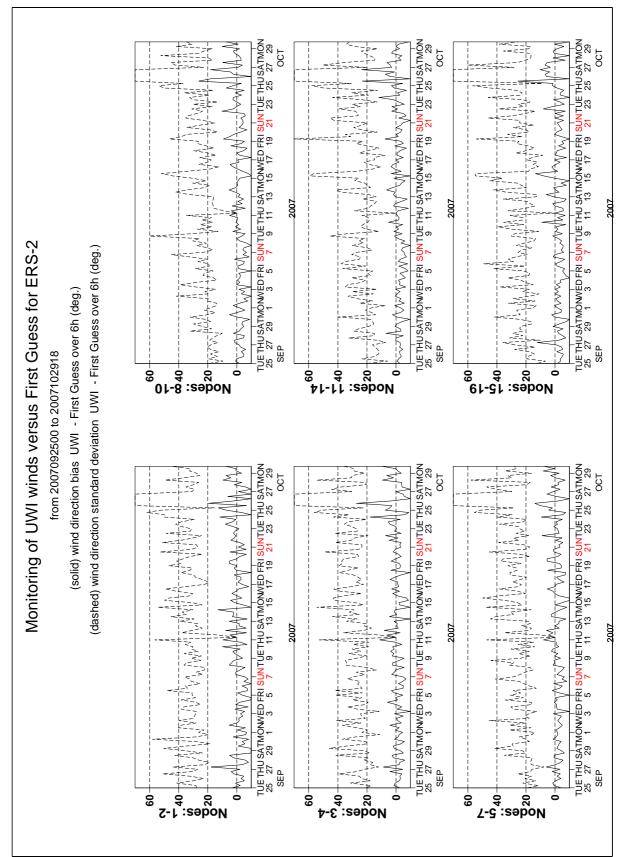


FIGURE 17 Same as Fig. 15, but for the wind direction difference. Statistics are computed only for wind speeds higher than 4 m/s.





ESA EOP-GQ

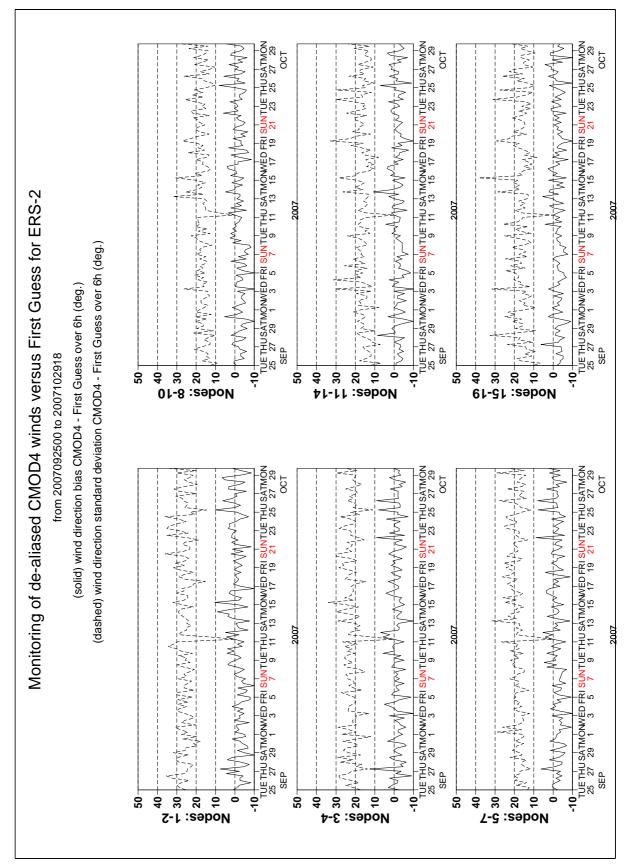
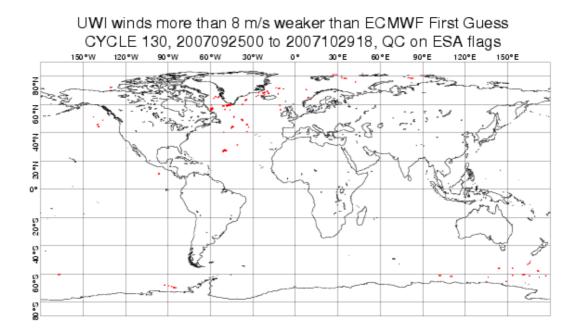


FIGURE 18 Same as Fig. 17, but for the de-aliased CMOD4 data.





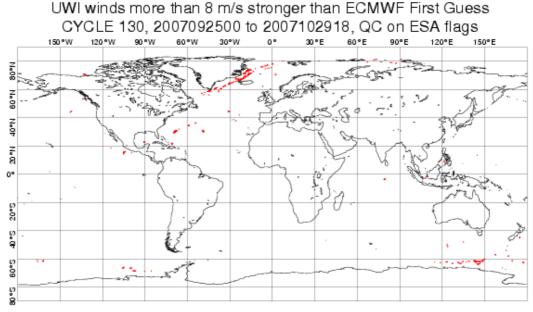


FIGURE 19 Locations of data during cycle 130 for which UWI winds are more than 8 m/s weaker (top panel) respectively stronger (lower panel) than FGAT, and on which QC on UWI flags and the ECMWF land/sea-ice mask was applied.



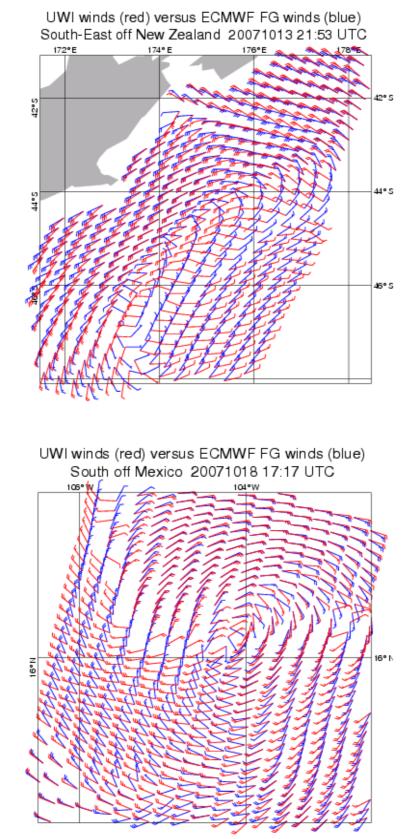


FIGURE 20 Comparison between UWI winds (in red) and ECMWF FG winds (in blue) for a case on 13 October 2007 south-east from New Zealand (top panel) and a case on 18 October south from Mexico (lower panel).



### 4.3.3 Scatter plots

Scatter plots of FG winds versus ERS-2 winds are displayed in Figures 21 to 24. Values of standard deviations and biases are slightly different from those displayed in Table 6. Reason for this is that, for plotting purposes, the in 0.5 m/s resolution ERS-2 winds have been slightly perturbed (increases scatter with 0.02 m/s), and that zero wind-speed ERS-2 winds have been excluded (decreases scatter with about 0.05 m/s).

The scatter plot of UWI wind speed versus FG (Figure 21) is very similar to that for (at ECMWF inverted) de-aliased CMOD4 winds (Figure 23). It confirms that the ESACA inversion scheme is working properly.

Winds derived on the basis of CMOD5 are displayed in Figure 24. The relative standard deviation is lower than for CMOD4 winds (1.44 m/s versus 1.48 m/s). Compared to ECMWF FG, CMOD5 winds are 0.41 m/s slower.



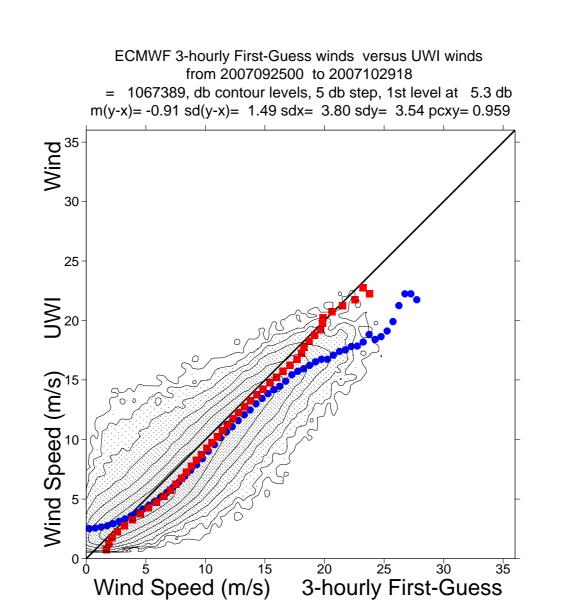


FIGURE 21 Two-dimensional histogram of first guess and UWI wind speeds, for the data kept by the UWI flags, and QC based on the ECMWF ice and land and sea-ice mask. Circles denote the mean values in the y-direction, and squares those in the x-direction.



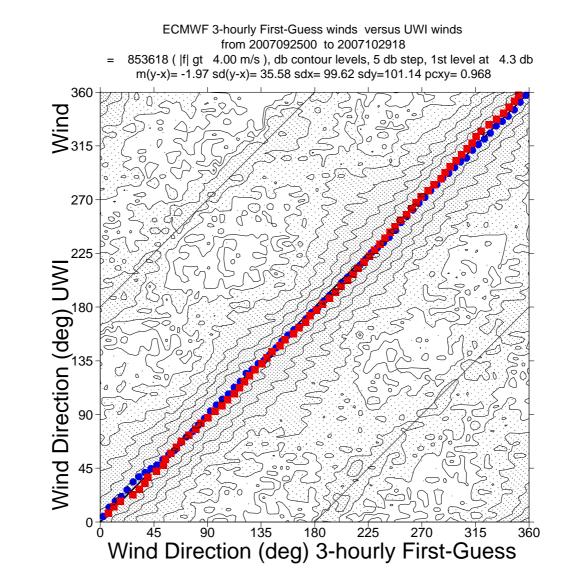


FIGURE 22 Same as Fig. 21, but for wind direction. Only wind speeds higher than 4m/s are taken into account.



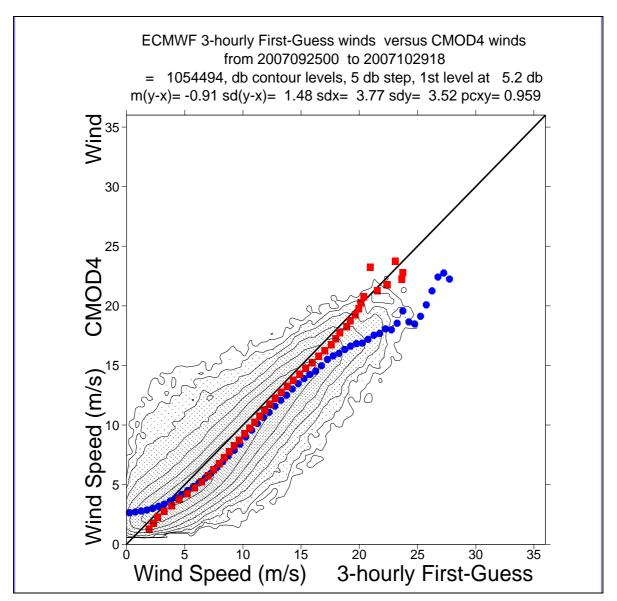


FIGURE 23 Same as Fig. 21, but for de-aliased CMOD4 winds.



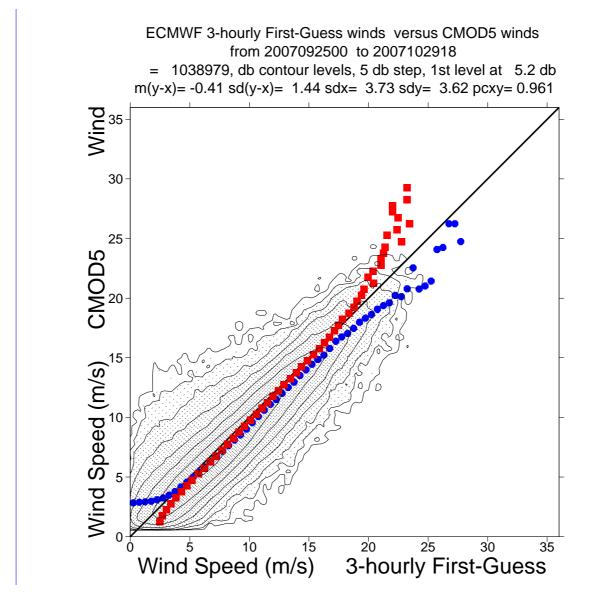


FIGURE 24 Same as Fig. 21, but for de-aliased CMOD5 winds.



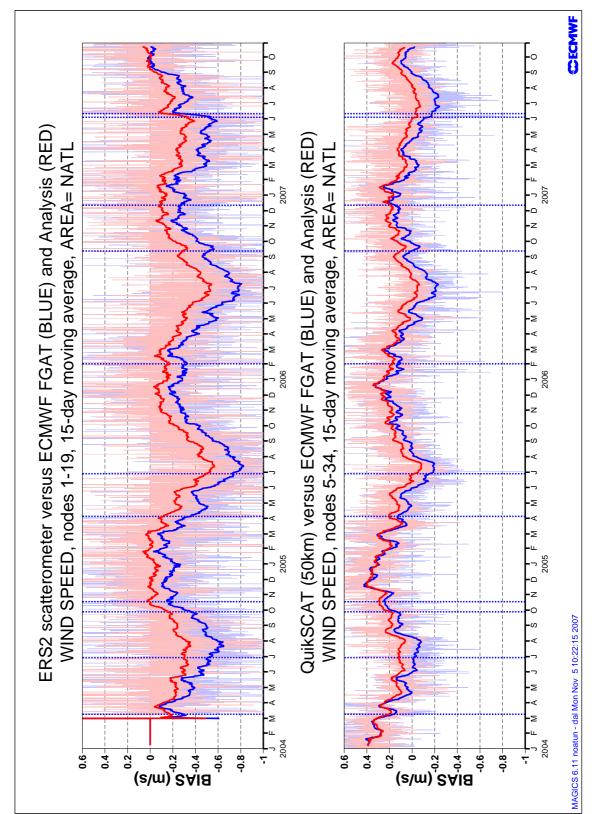


FIGURE 25 Bias relative to FG winds for actively assimilated ERS-2 winds (based on CMOD5) for nodes 1-19 (top panel) respectively of 50-km QuikSCAT (based on the QSCAT-1 model function and reduced by 4%) for nodes 5-34 (lower panels) averaged over the area (20N-90N, 80W-20E), and displayed for the period 01 January 2004 – 29 October 2007. Fat curves represent centered 15-day running means, thin curves values for 6-hourly period. Vertical dashed blue lines mark ECMWF model changes



### 4.3.4 Timeliness evolution

The Scatterometer product timeliness is defined as the difference between the acquisition time of the first product and the creation date of the file received in ESRIN-PCS. Once the UWI file is received in ESRIN, data are converted in BUFR format and sent to users via the GTS network. Therefore that timeliness is an indicator of the delay time that the user could expect in the data dissemination. The analysis does not take into account delays in the GTS network. For each file received from the ground station, the timeliness is computed and this analysis reports the daily mean timelines obtained by averaging all the values.

The analysis has been performed on the daily timeliness average. Timeliness is zero when no products are received.

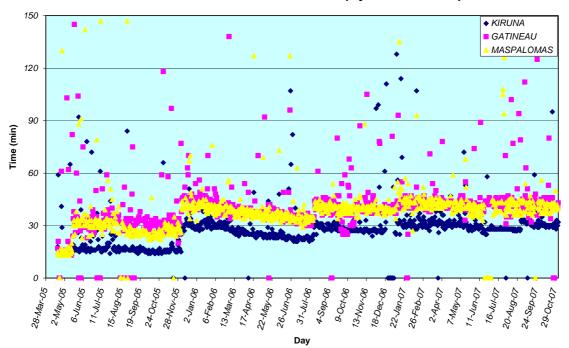
In the next figures is showed the evolution of the daily mean timeliness of Kiruna, Maspalomas, Gatineau, West Freugh, Miami, McMurdo and Beijing products until 29<sup>th</sup> October 2007 (end of Cycle 130). The starting date of the analysis, for each station, is reported in the following table:

STATION	START DATE	
Kiruna	19 April 2005	
Gatineau	19 April 2005	
Maspalomas	19 April 2005	
West Freugh	19 April 2005	
Miami	19 April 2005	
McMurdo	13 March 2007	
Beijing	13 March 2007	

 TABLE 6 Starting date of Timeliness analysis for each station

The Figure 26 shows the results of the investigation for Gatineau, Kiruna and Maspalomas stations.





**TIMELINESS EVOLUTION (Cycles 105 - 130)** 

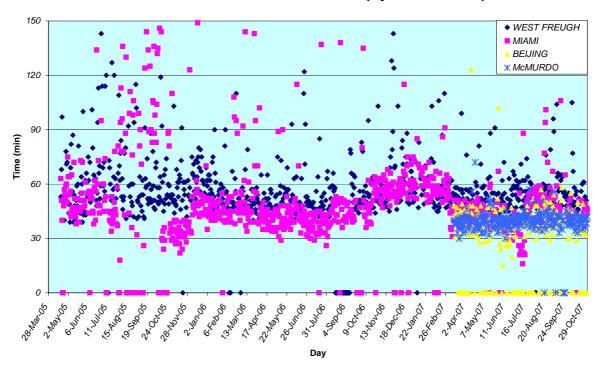
FIGURE 26 : Timeliness evolution from 19 April 2005 (mid cycle 105) to 29<sup>th</sup> October 2007 (end of cycle 130) for Gatineau, Kiruna and Maspalomas ground stations (Timeliness is zero when no products are received).

Apart from some values out of the general tendency due to temporary system or connection problem, since the beginning of the analyzed period a timeliness increase is detected for all the stations. In particular, it can be recognized a discontinuous trend for the three stations with quickly increases in the same days for the 3 stations followed by a slightly decrease in the subsequent months. In depth analysis showed that these rapid increases occur about in the following days: 5 May 2005, 5 December 2005, 9 August 2006 and 9 January 2007. This behavior could be due to settings modifications in the ground segment.

Nowadays Kiruna products have the lowest timeliness with an average value of 30 minutes. Gatineau and Maspalomas have timeliness around 45 minutes.

The analysis for West Freugh, Miami, Beijing and McMurdo stations is showed in Figure 27.





**TIMELINESS EVOLUTION (Cycles 105 - 130)** 

FIGURE 27 : Timeliness evolution from 19 April 2005 (mid Cycle 105) to 29<sup>th</sup> October 2007 (end of Cycle 130) for West Freugh, Miami, Beijing and McMurdo ground stations (Timeliness is zero when no products are received).

West Freugh and Miami stations show a similar regular trend in the analyzed period. More in detail a slightly increased timeliness could be identified since October 2006 followed by a decrease since January 2007. In the last cycle West Freugh products had an average timeliness of 50 minutes, Miami around 40 minutes. The analysis for Beijing and McMurdo stations shows a lower timeliness that is around 40 minutes for both stations.

The analysis carried out shows that till December 5th 2005 UWI products delivered from the three ESA ground station (Kiruna, Maspalomas, Gatineau) had a timeliness that fulfils the requirements for nowcasting application (data received on average within 25 minutes). After that date performances degraded and nowadays the average timeliness is around 35 minutes. This trend needs further investigations to better understand the cause.



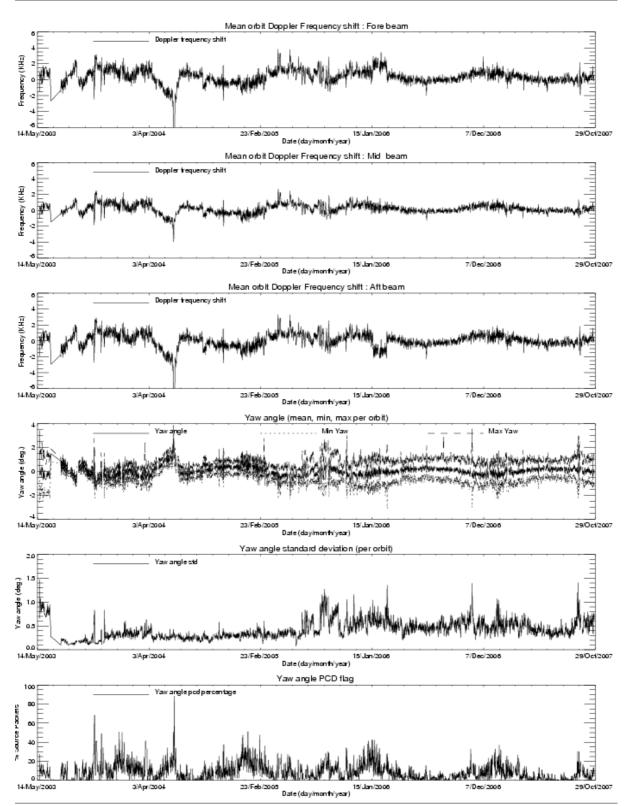
# 5 Yaw error angle estimation

The yaw error angle estimation is computed on-ground by the ESACA processors. The full set of results of the yaw processing is stored in an internal ESA product named HEY (Helpful ESA Yaw) disseminated from the ground station to ESRIN. The estimation of the yaw error angle is based on the Doppler shift measured on the received echo. That estimation can be done with a good accuracy only for small yaw error angle (in the range between +/-4 deg.). Above that range, due to high Doppler frequency shift the signal spectrum is outside the receiver bandwidth and the yaw estimation is strong degraded. Details regarding the yaw be found following document processing can on the (chapter 9): http://earth.esa.int/pcs/ers/scatt/articles/soamain-030521.pdf.

The yaw error angle estimation aims to compute the correct acquisition geometry for the three Scatterometer antenna throughout the entire orbit. The Yaw error angle information is used in the radar equation to derive the calibrated backscattering (sigma nought) from the Earth surface and to select the echo samples associated to one node. In ESACA the definition of the node position is as the one adopted in the old processor (for details see:.http://earth.esa.int/pcs/ers/scatt/articles/scatt\_work98\_processing.pdf). In such way the distance between the nodes (both along and across track) is kept constant (25 Km) and what is changing in function of the yaw error angle is the number of echo samples that contributes to the node calculation and the incidence angle of the measurement. This because the three Scatterometer antennae could see the node with a different geometry due to an arbitrary variation of the yaw angle along track. The number of samples that actually contributes to a node and the yaw flag can be retrieved from the UWI Data Set Record (DSR) product. For that reason the definition of few fields in the UWI product has been updated. For details see the Scatterometer cyclic report - cycle 90 -. The Figure 28 (since beginning of HEY dissemination) and Figure 29 (cycle) show for each orbit the average Doppler frequency shift (first 3 plots Fore Mid and Aft antenna), the minimum, maximum and mean yaw (fourth plot), the yaw standard deviation (fifth plot) and the percentage of source packets acquired with a yaw error angle outside the range +/-2 degrees (sixth plot). On average the yaw evolution is within the specification for the ESACA processor to assure calibrated data. The evolving yaw bias occurred in June 2004 has been reported to the flight segment and corrective actions have been put in place to compensate for.

The result of the monitoring for cycle 130 is an average (per orbit) yaw error angle within the expected nominal range (+/- 2 degrees) for most of the orbit. Peak in the yaw angle standard deviation on  $22^{nd}$  October could be due to the satellite attitude not corrected caused by yaw statistics not disseminated to Esoc caused by ground segment network problems





#### ERS-2 WindScatterometer: DOPPLER & YAW (HEY)

FIGURE 28 Doppler frequency shift and Yaw error angle evolution since August 2003 with a smooth of 14 orbits



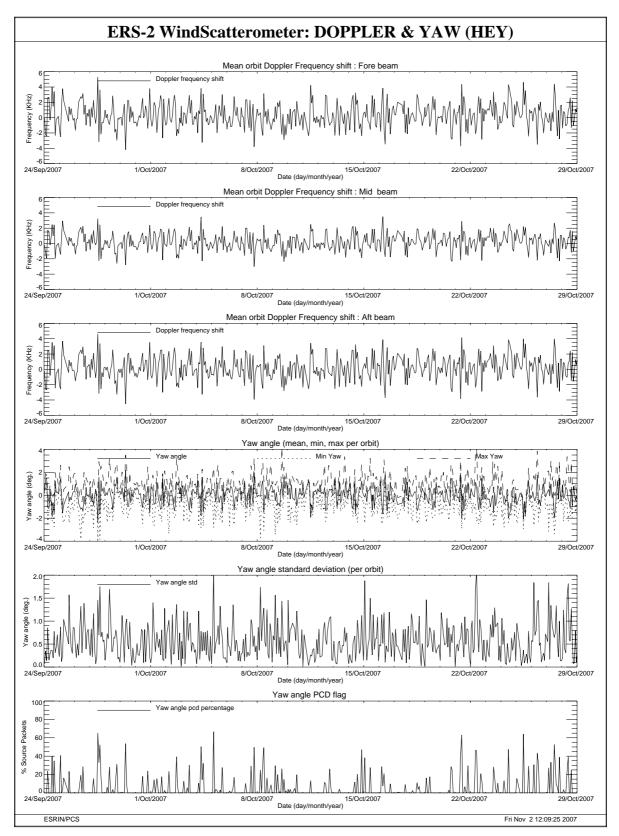


FIGURE 29 Doppler frequency shift and Yaw error angle evolution cycle 130.

